

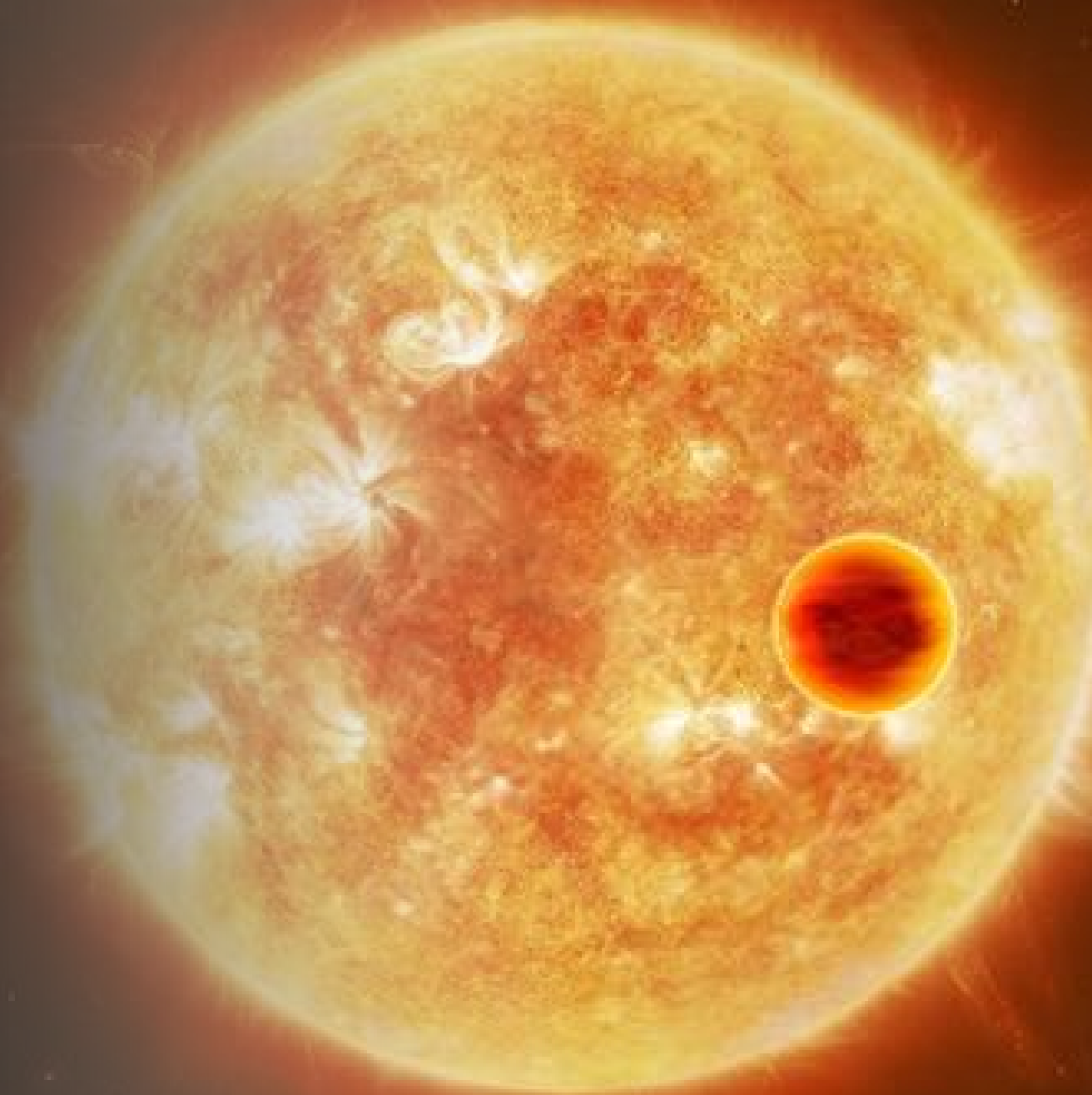


Exploring Rocky Worlds: On the Precipice of a New Frontier

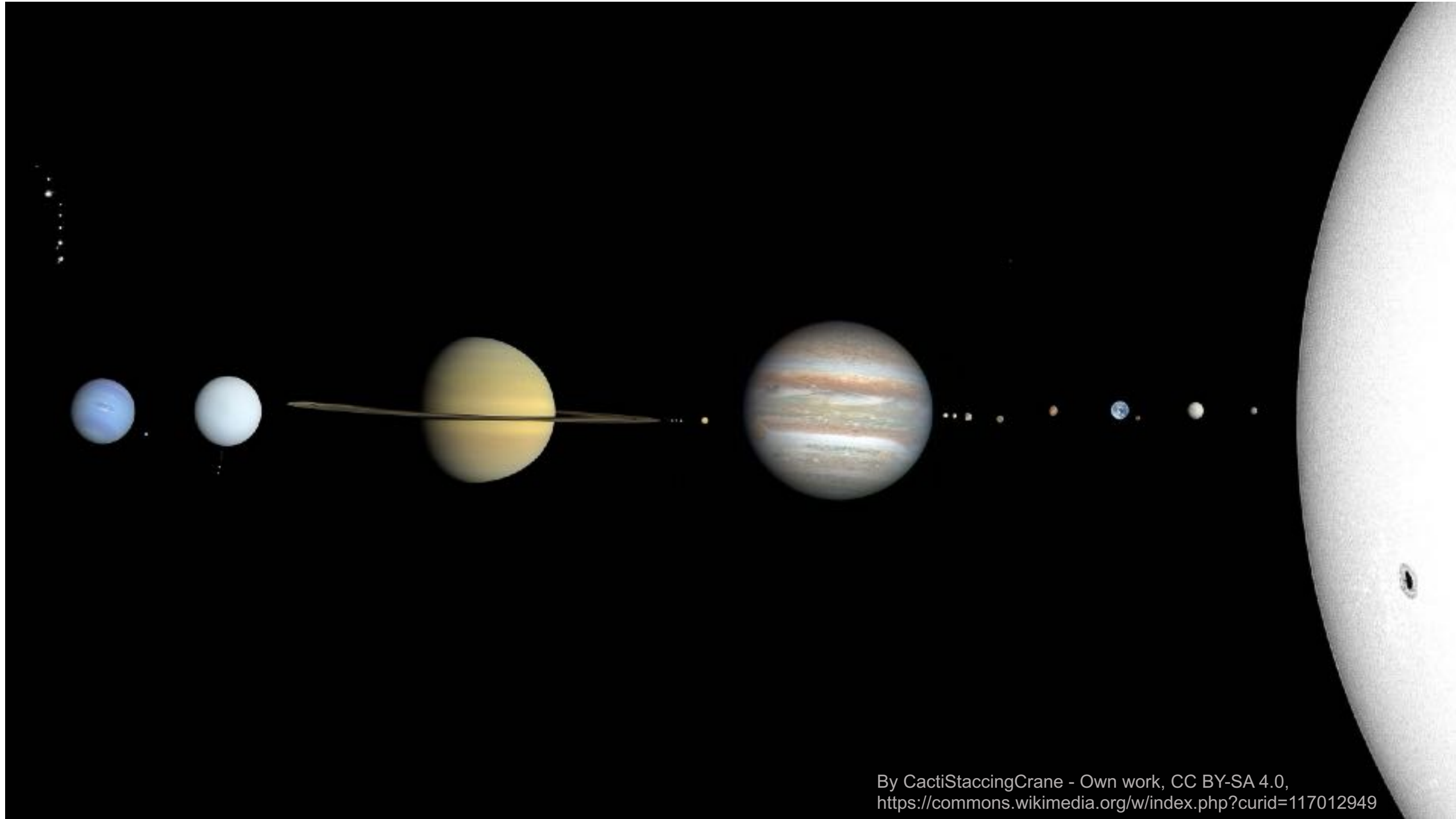
Katie Bennett

Ph.D. Student

Johns Hopkins University







You are here



You are here



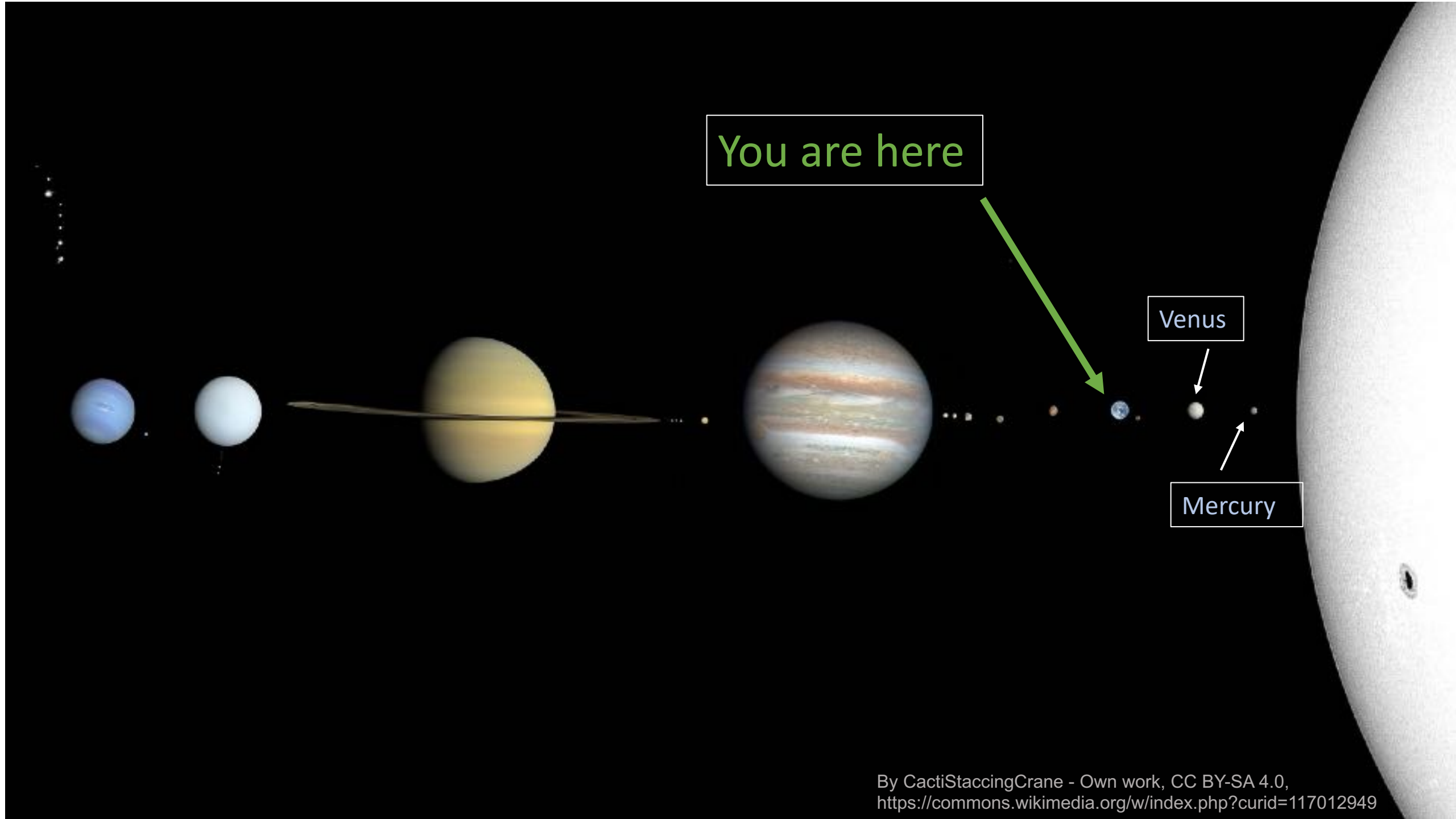
Mercury



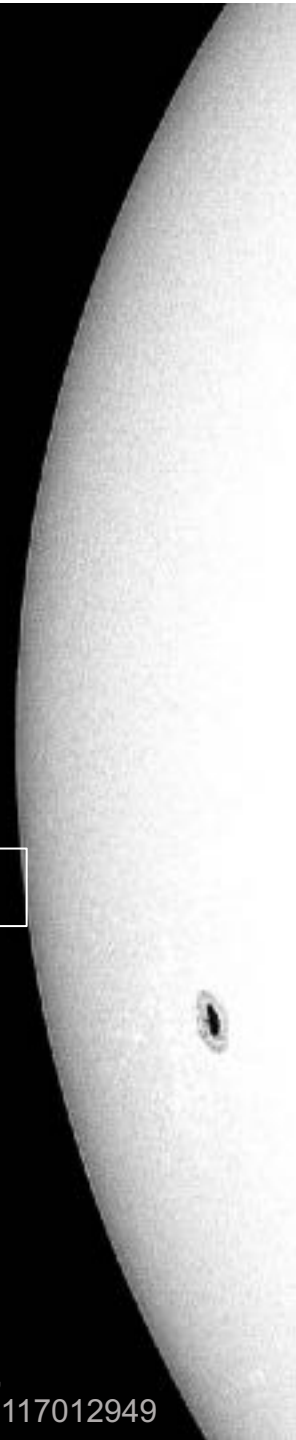
You are here

Venus

Mercury



You are here



Mars

Venus

Mercury

You are here



Venus

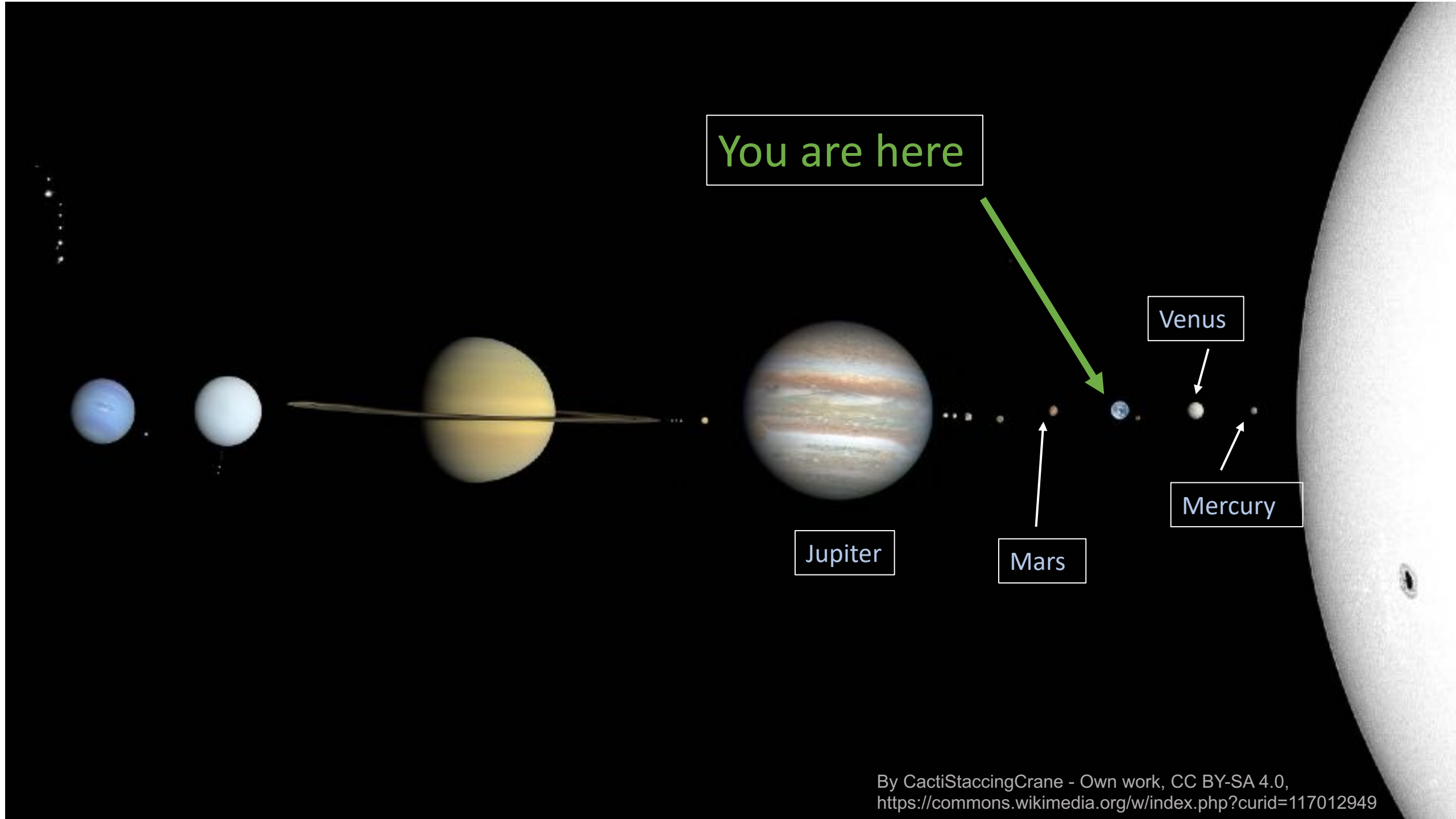


Mercury



Jupiter

Mars



You are here



Saturn

Jupiter

Mars

Venus

Mercury

You are here



Uranus

Saturn

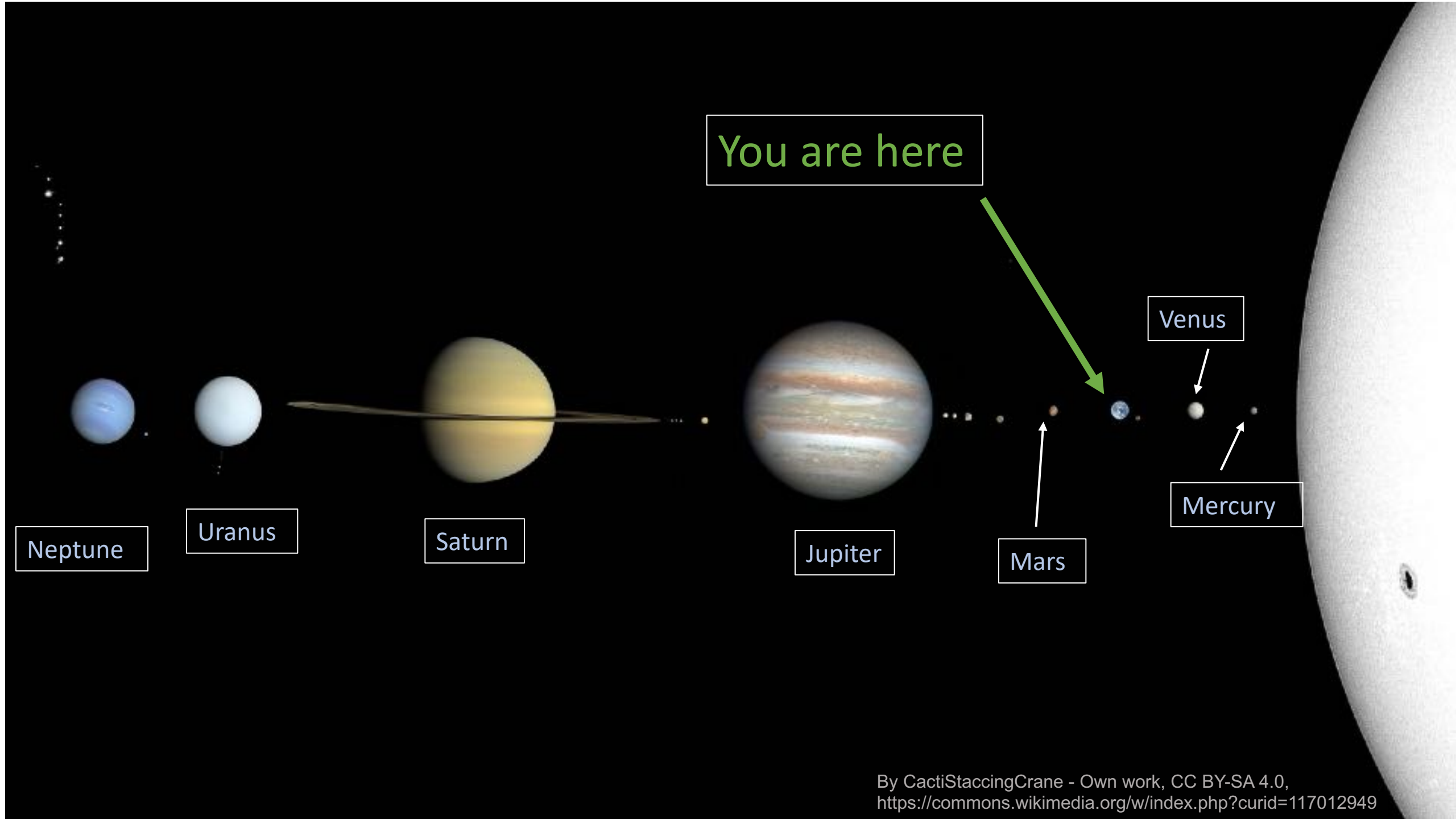
Jupiter

Mars

Venus

Mercury





You are here

Neptune

Uranus

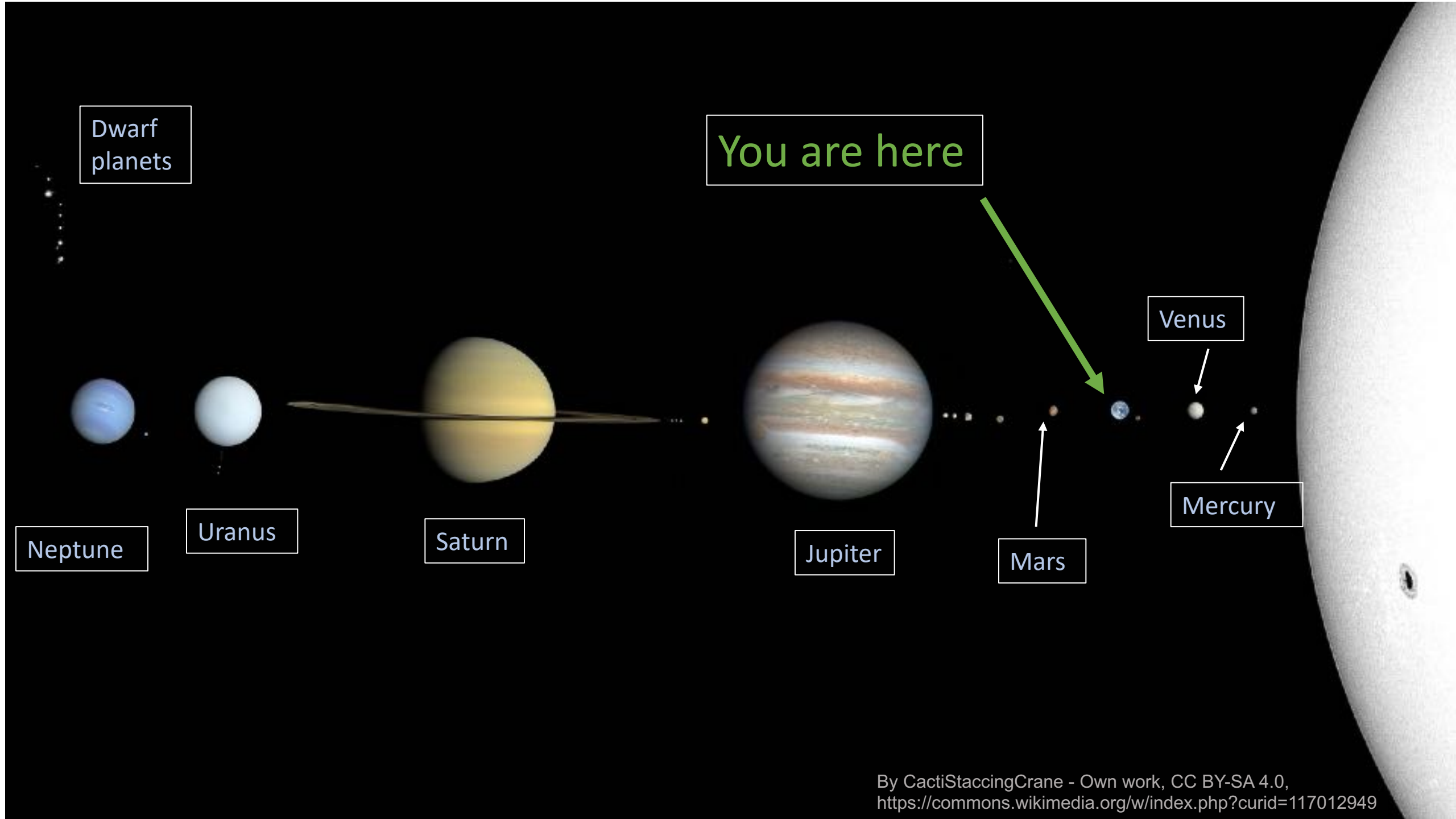
Saturn

Jupiter

Mars

Venus

Mercury



Dwarf planets

You are here

Neptune

Uranus

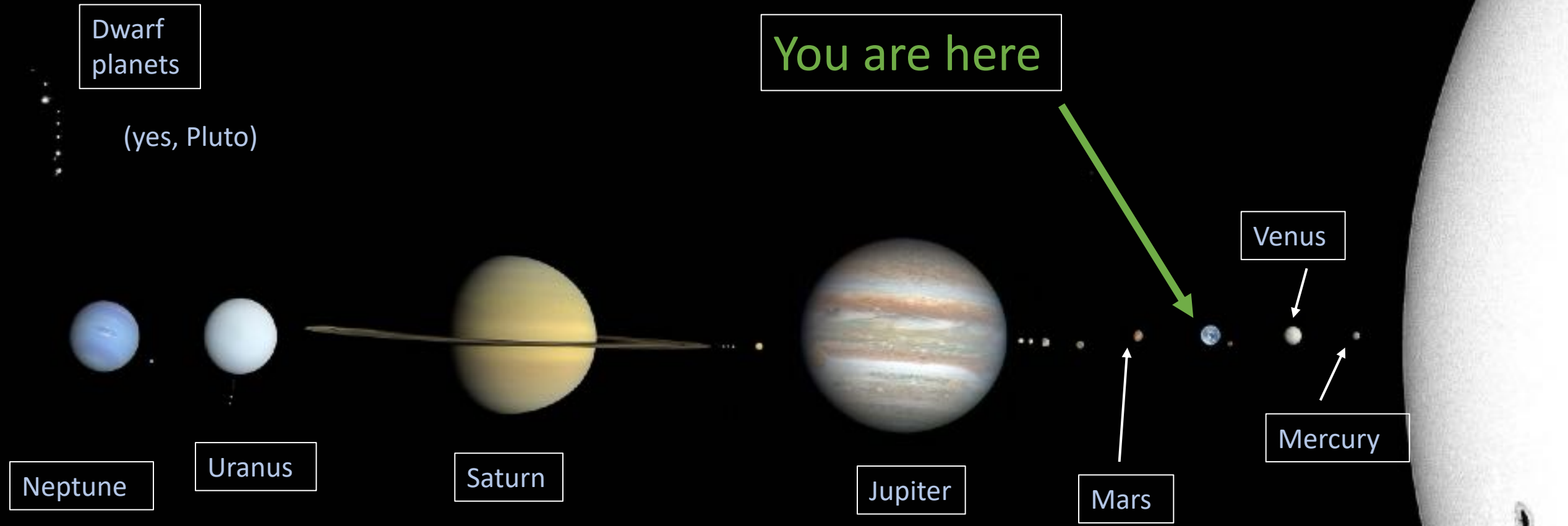
Saturn

Jupiter

Mars

Venus

Mercury



Dwarf planets

(yes, Pluto)

You are here

Neptune

Uranus

Saturn

Jupiter

Mars

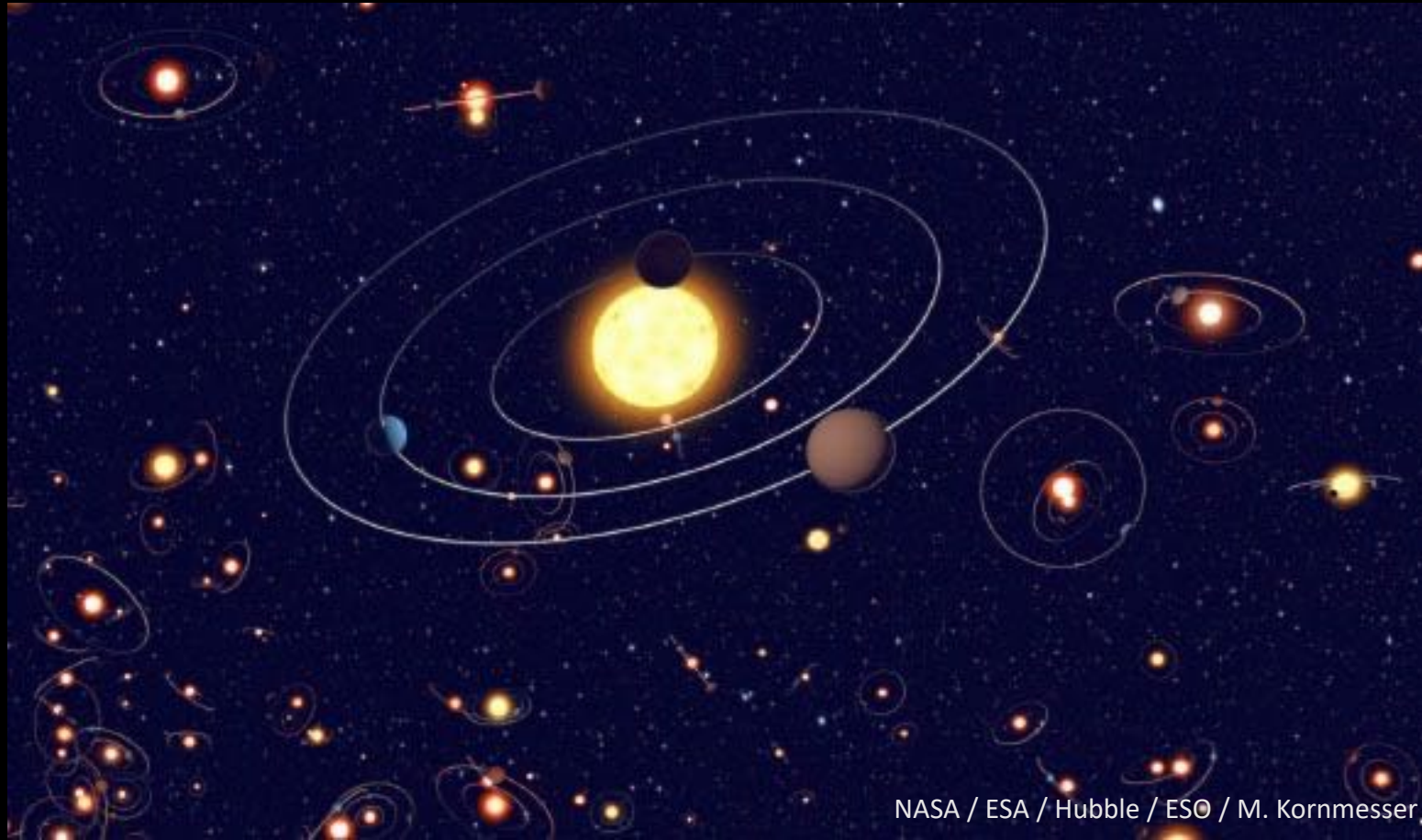
Venus

Mercury

YOU ARE HERE

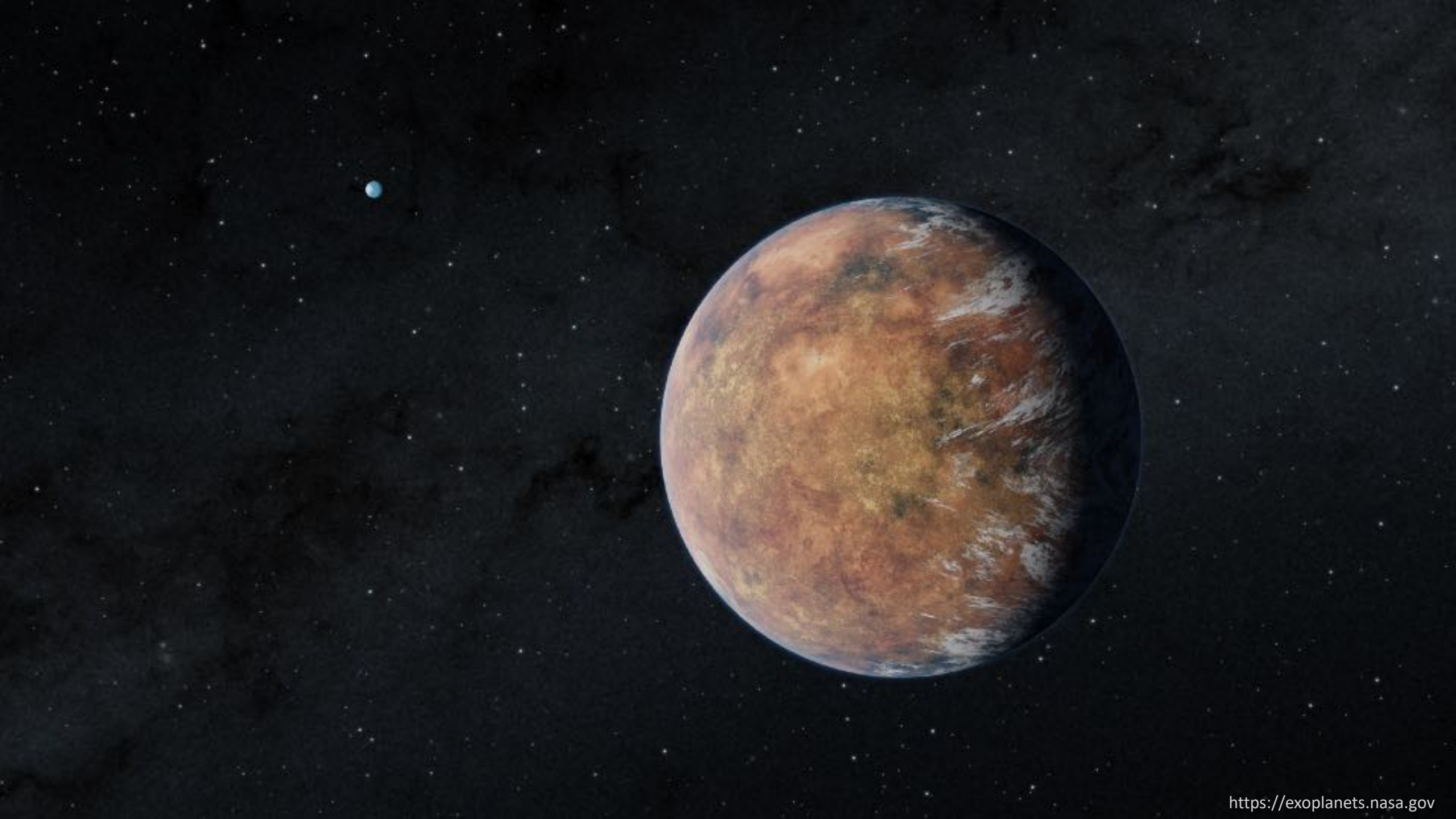




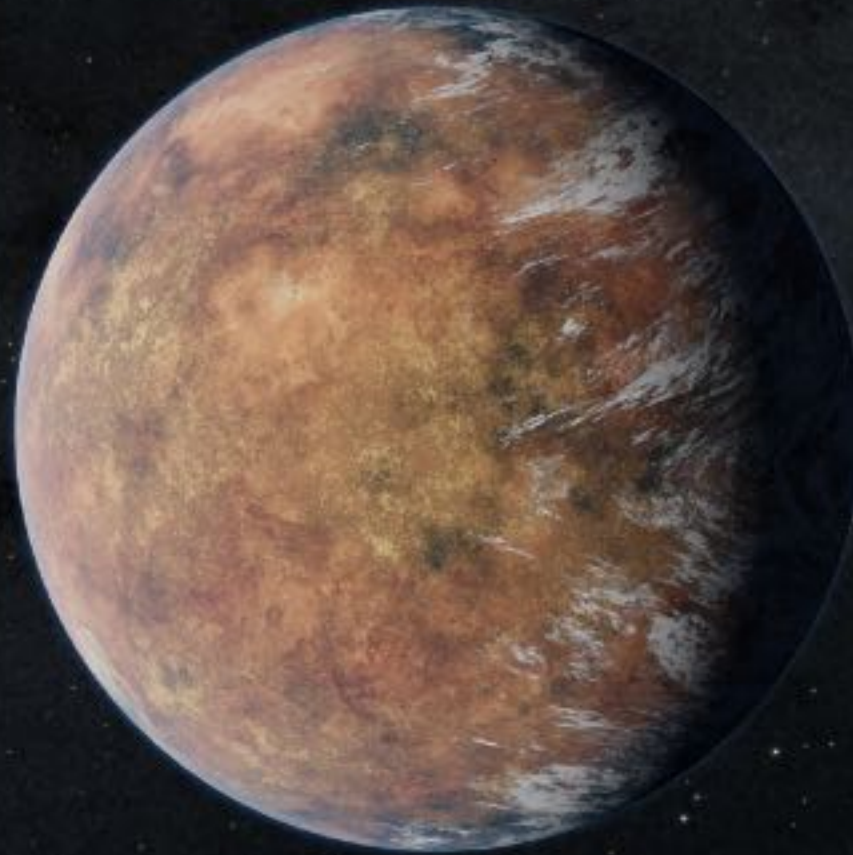


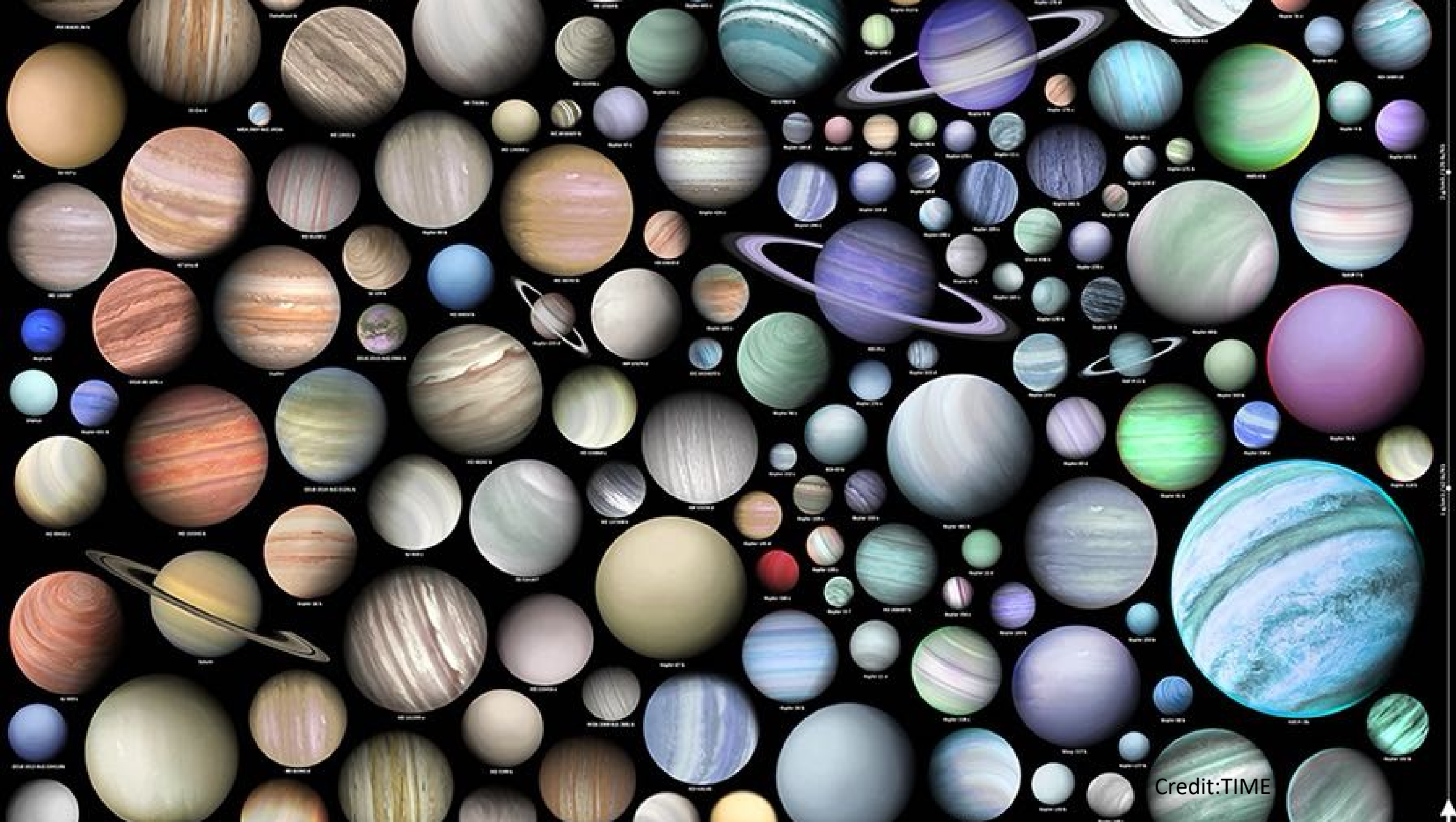
NASA / ESA / Hubble / ESO / M. Kornmesser

There are **MANY** other solar systems out there!

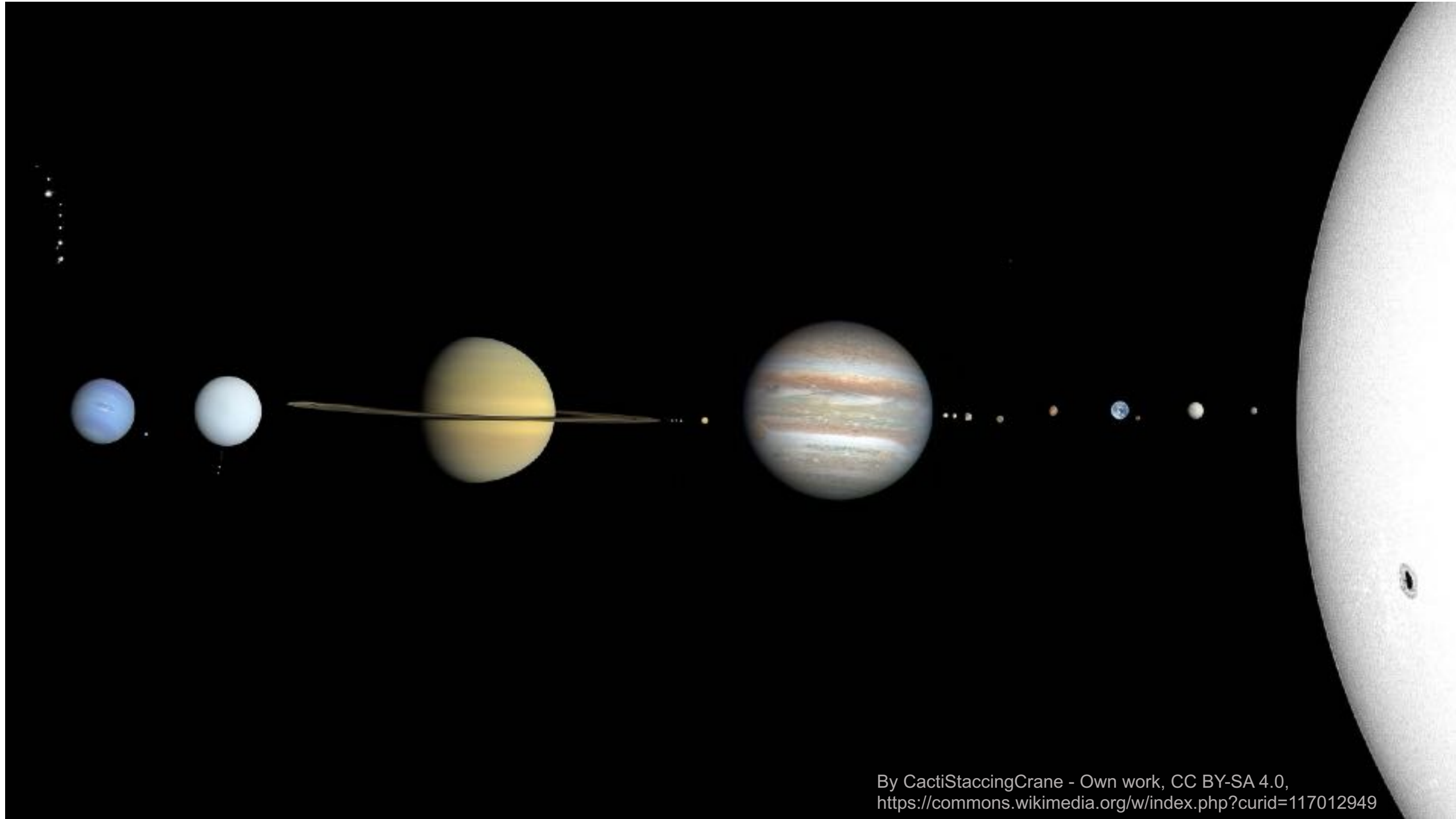


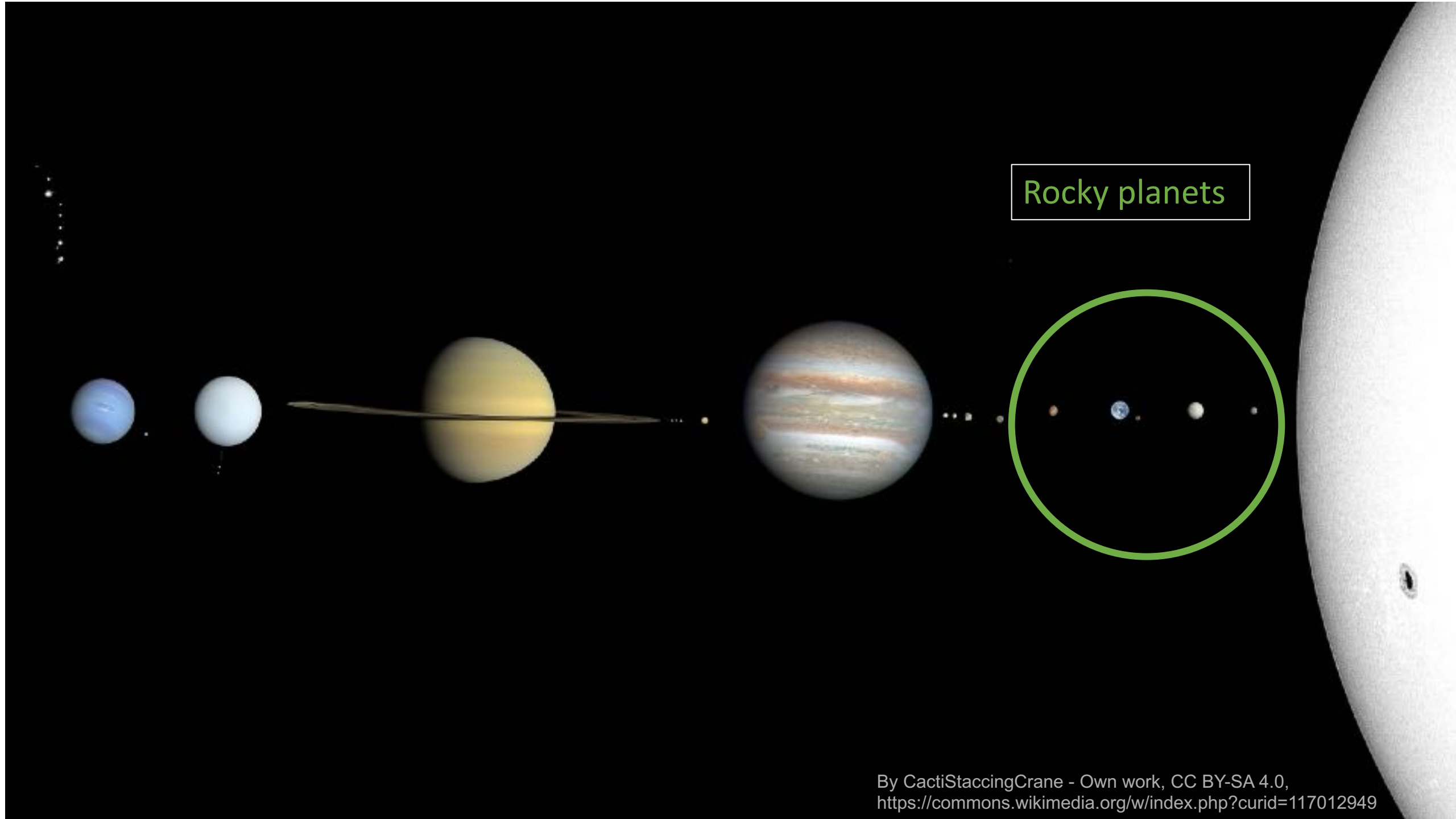
Exoplanet: a planet
orbiting a star other than
the Sun





Credit:TIME

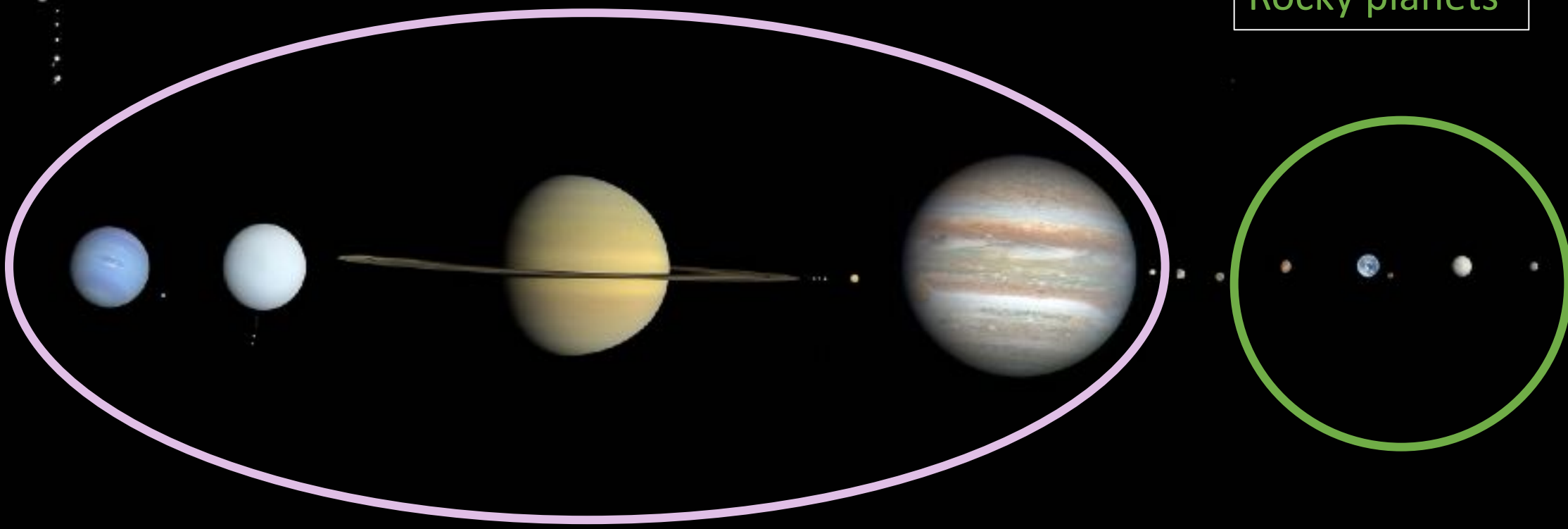




Rocky planets

Gas/ice giants

Rocky planets



Exoplanet Types



Gas Giants

The size of Saturn or Jupiter, or much larger. They include "hot Jupiters"- scorching planets in close orbits around their stars.



Neptune-Like

Similar in size to our own Neptune and Uranus, with hydrogen or helium-dominated atmospheres. "Mini-Neptunes," not found in our solar system, are smaller than Neptune but larger than Earth.



Terrestrial

Earth-sized or smaller, mostly made of rock and metal. Some could possess oceans or atmospheres and perhaps other signs of habitability.



Super-Earth

Typically "terrestrial," or rocky, and more massive than Earth but lighter than Neptune. They might or might not have atmospheres.

Are there
exoplanets that
are like Earth?







With JWST, we are at the precipice of a new frontier.

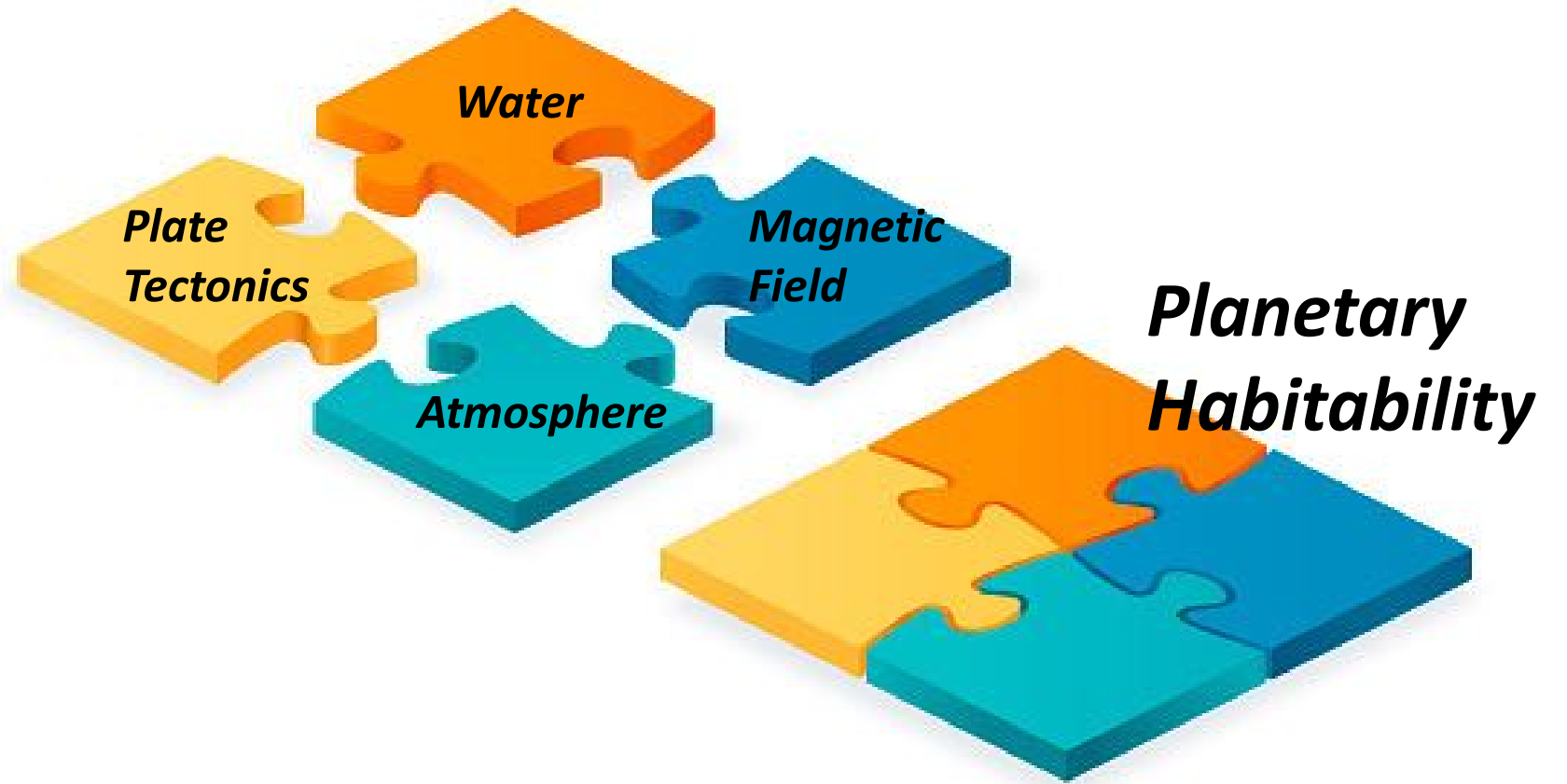
We can (try to) look at the atmospheres of rocky exoplanets for the first time!

*Why do we care about
planet atmospheres?*

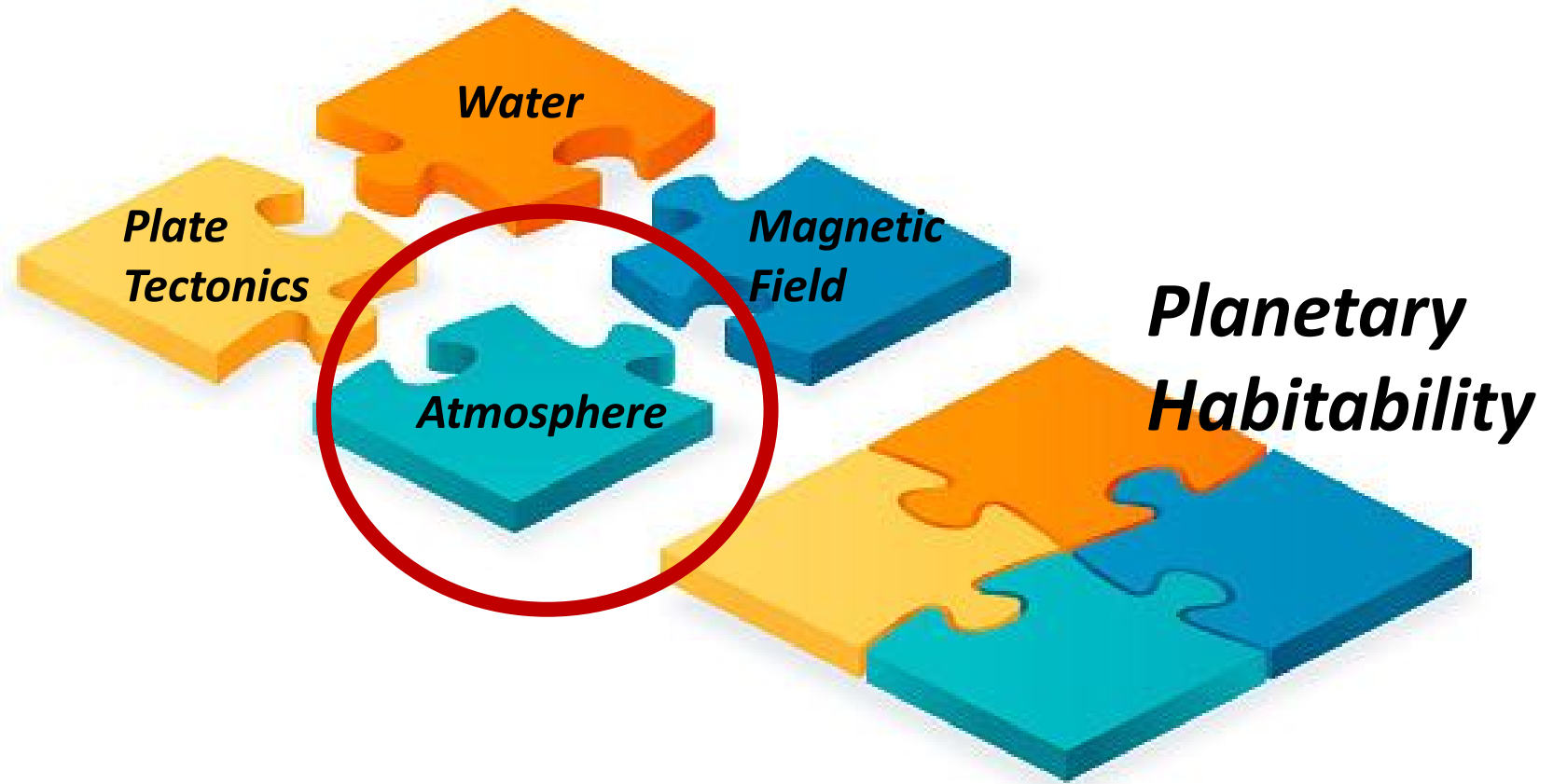
Are there
exoplanets that
are like Earth?



**Many different
components that
make up
planetary
habitability**



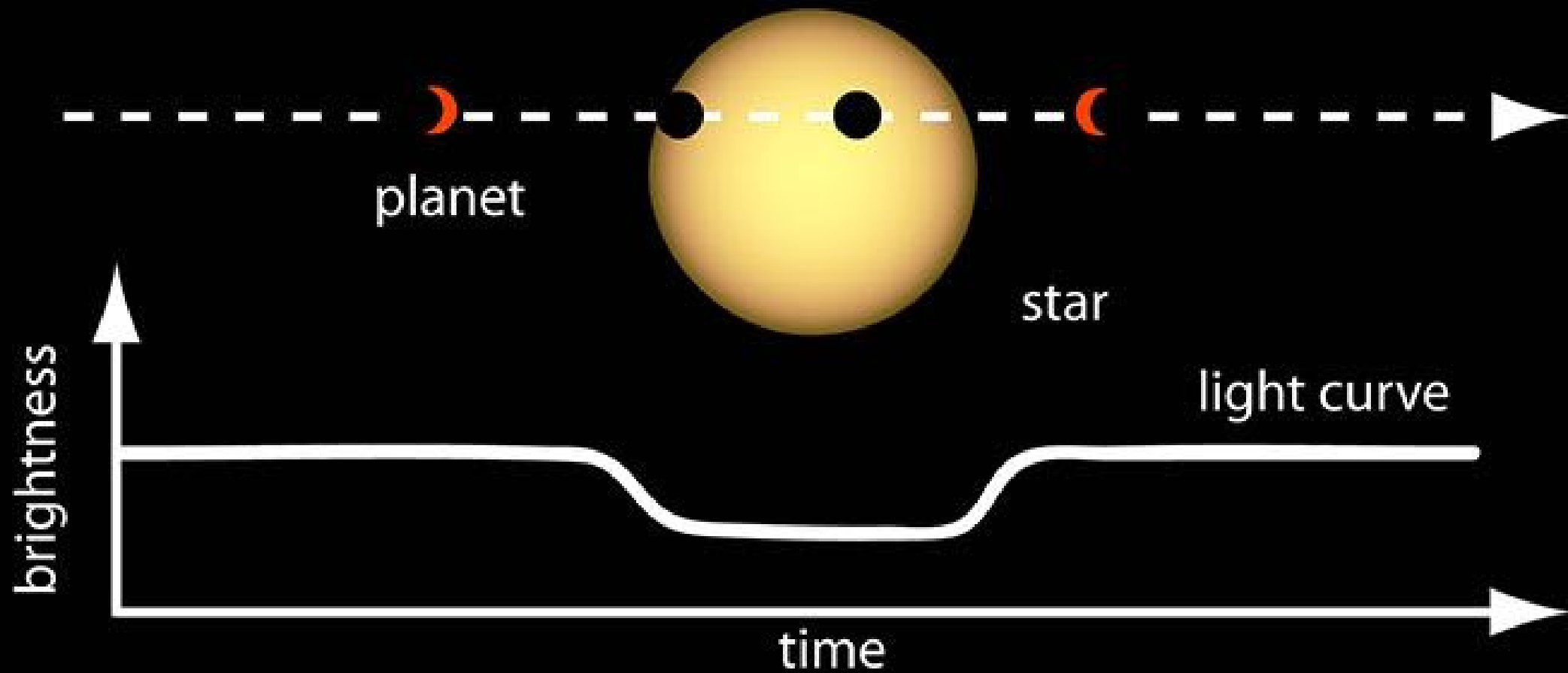
**Many different
components that
make up
planetary
habitability**

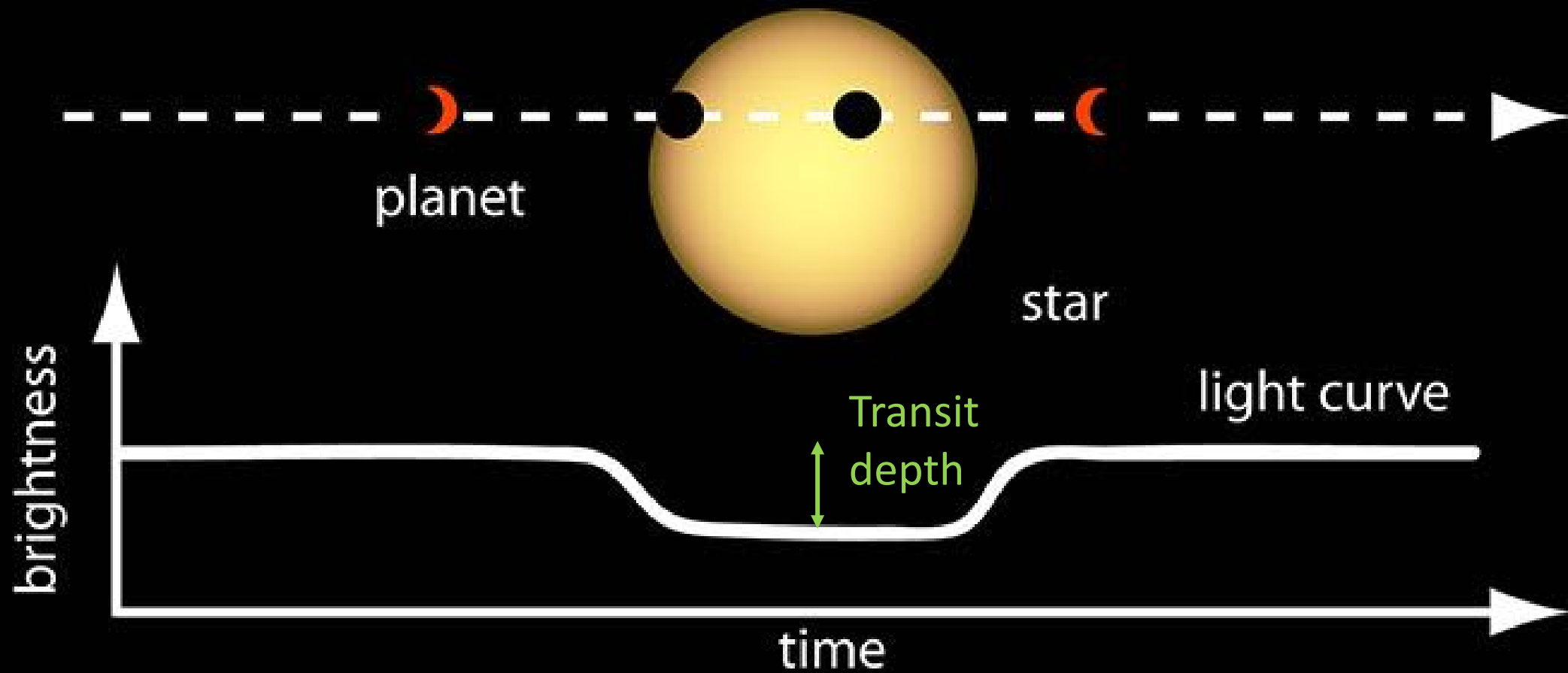




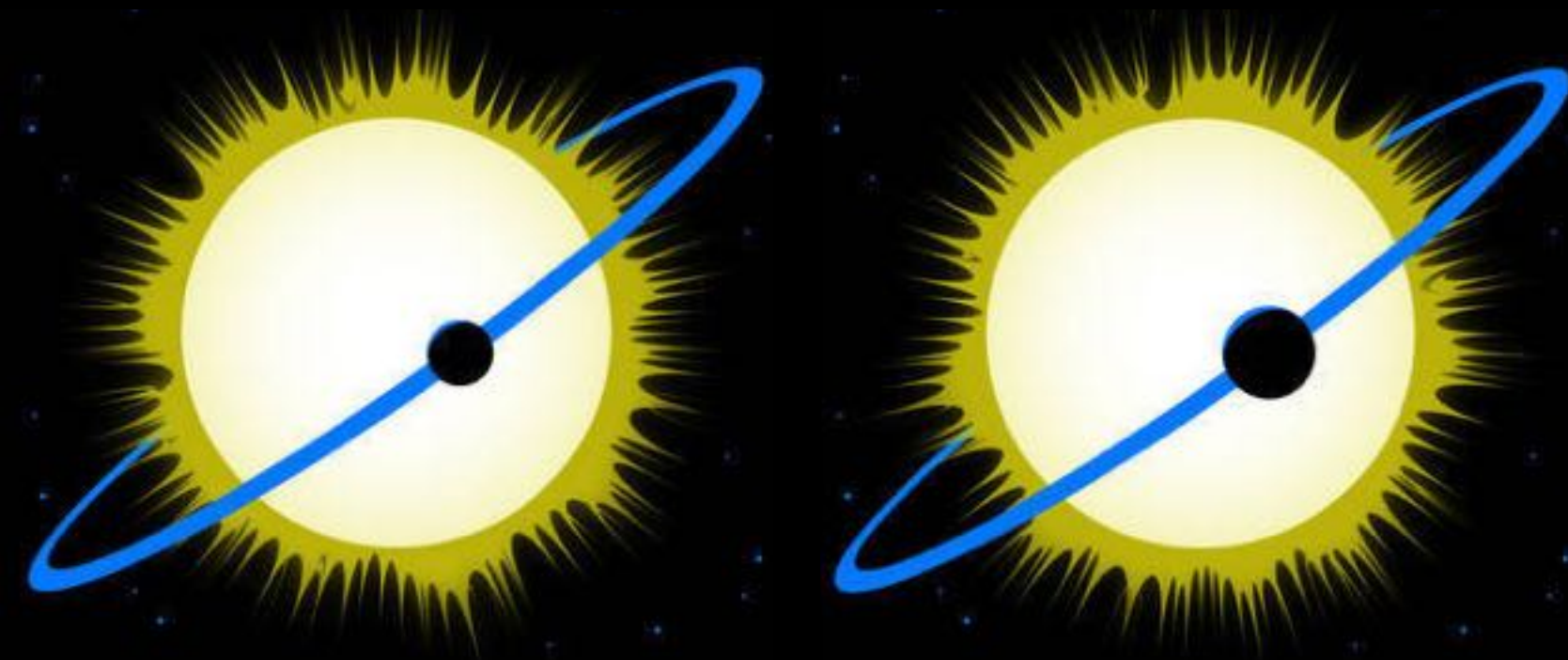
**HOW DO
YOU DO
THIS?**



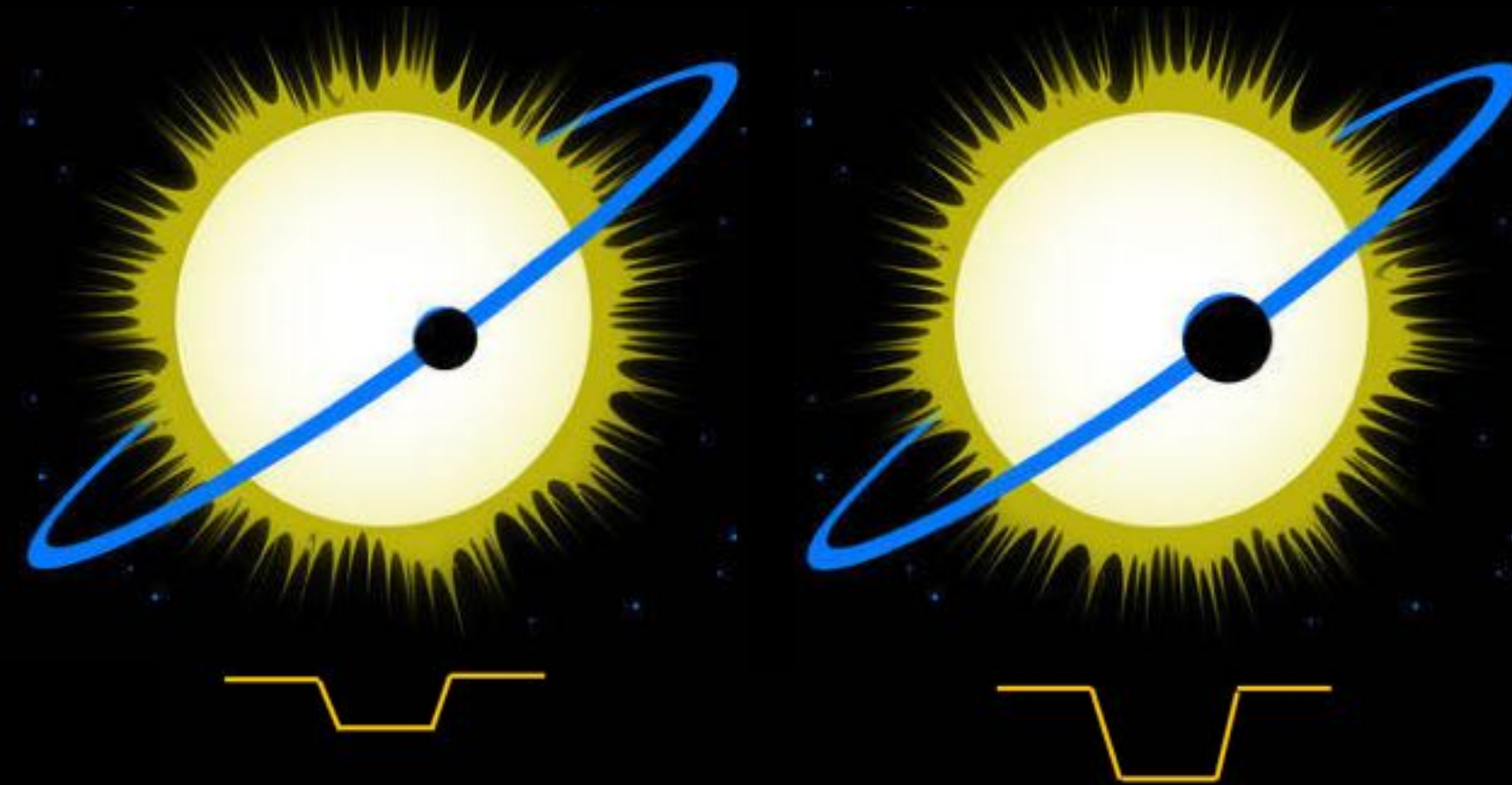




The size of the planet AND star matter!

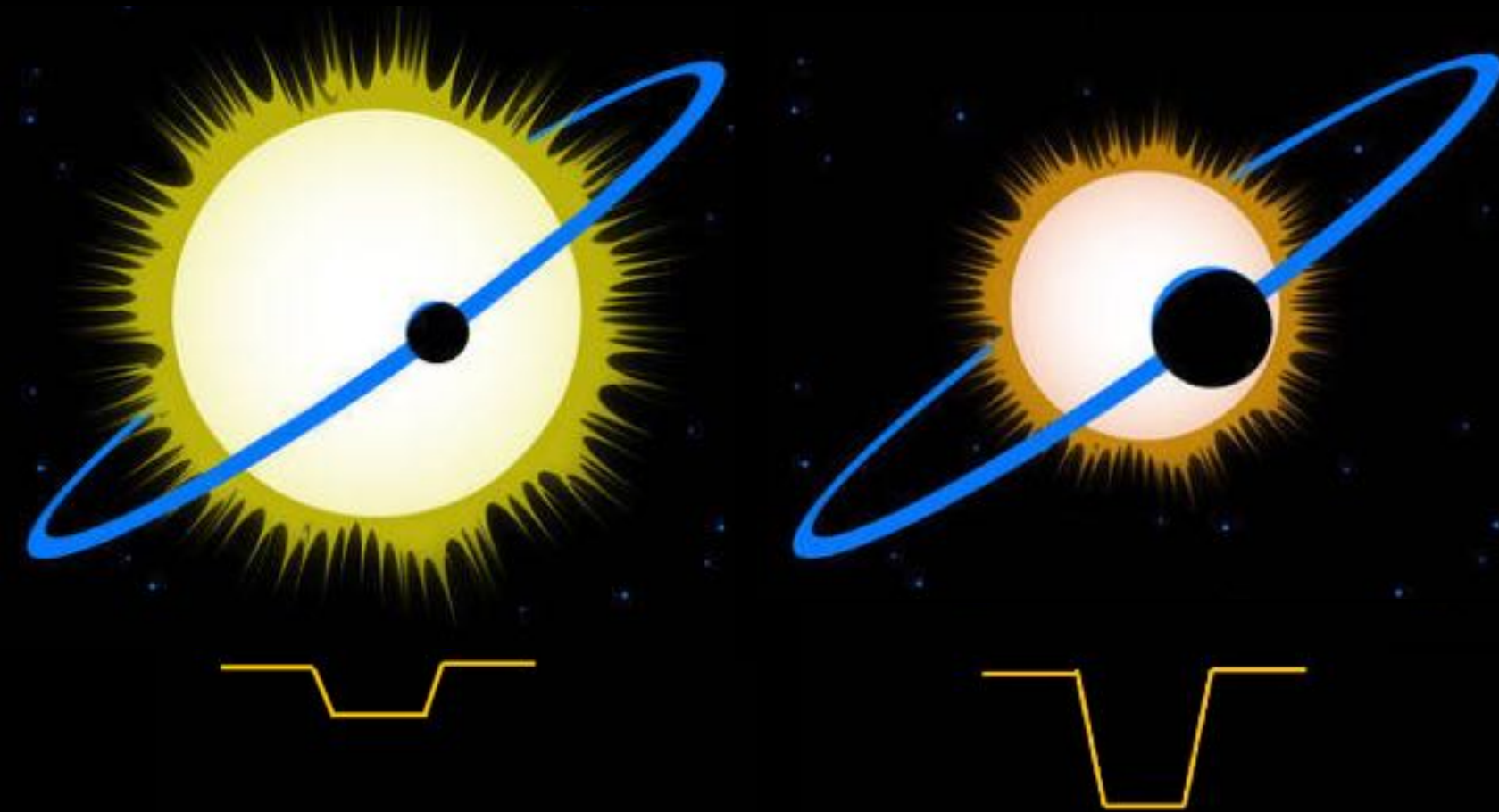


The size of the planet AND star matter!

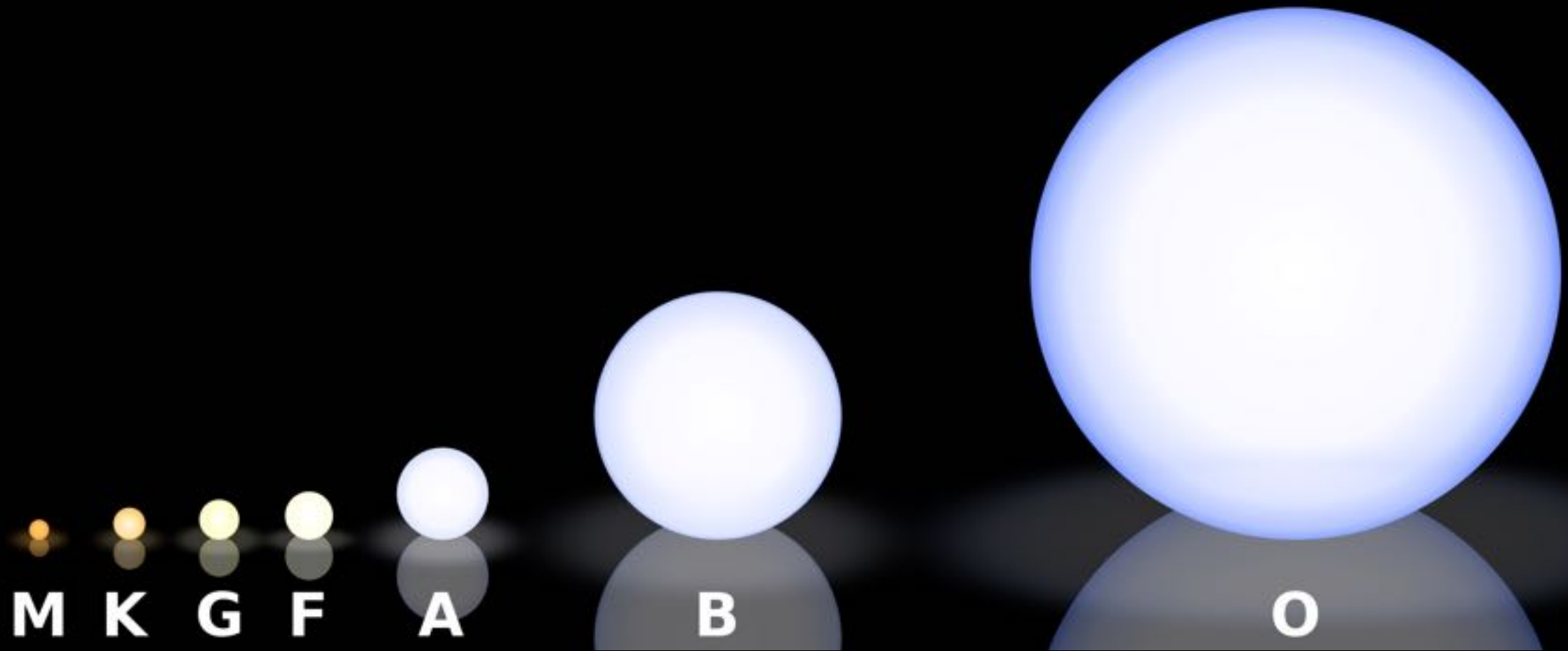


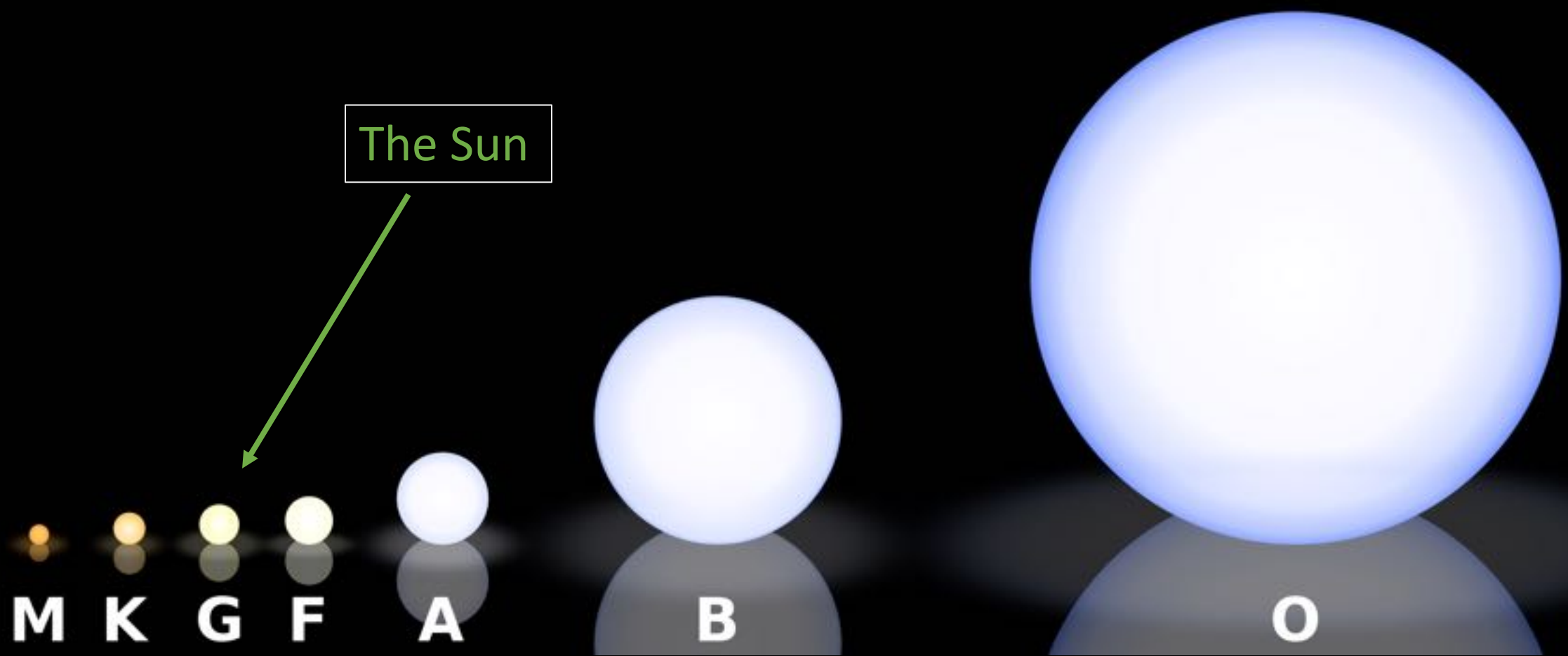
Easier to detect
BIGGER planets

The size of the planet AND star matter!



Easier to detect
BIGGER planets
around SMALLER
stars





The Sun

M

K

G

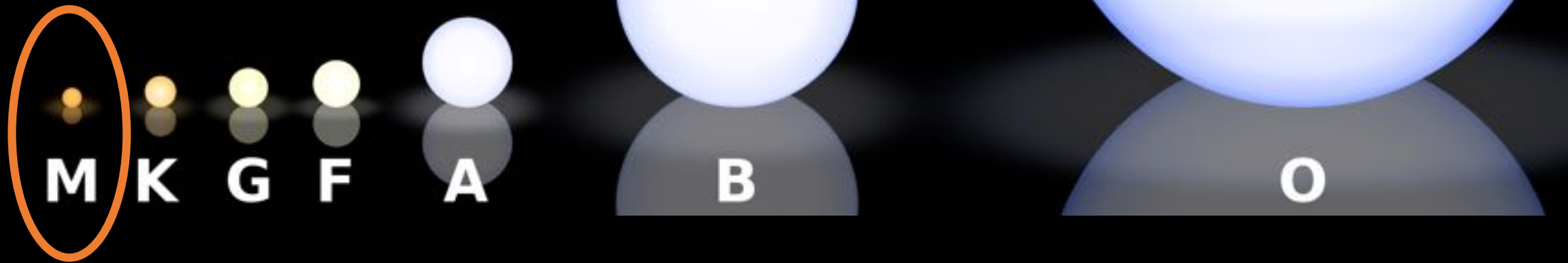
F

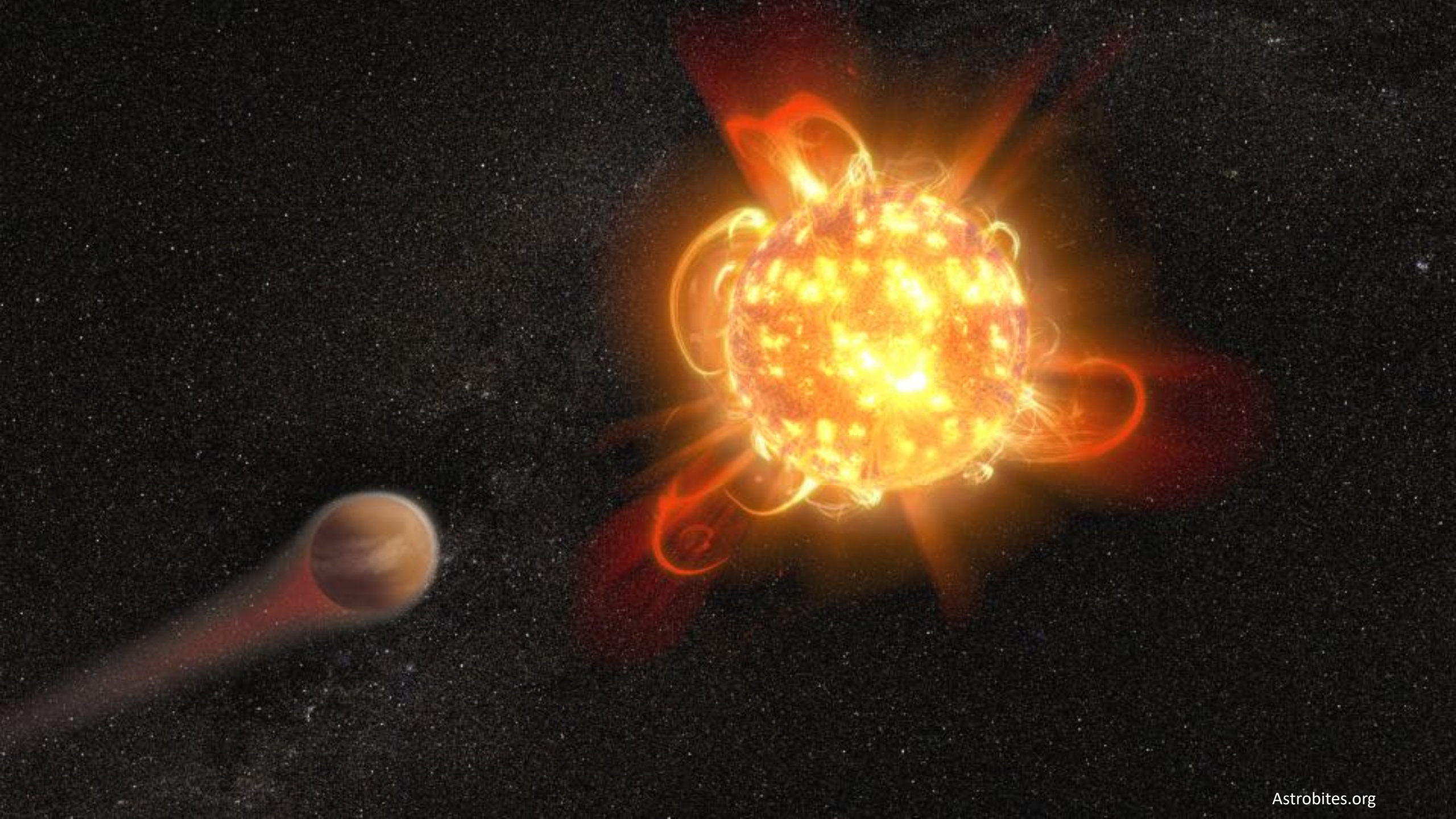
A

B

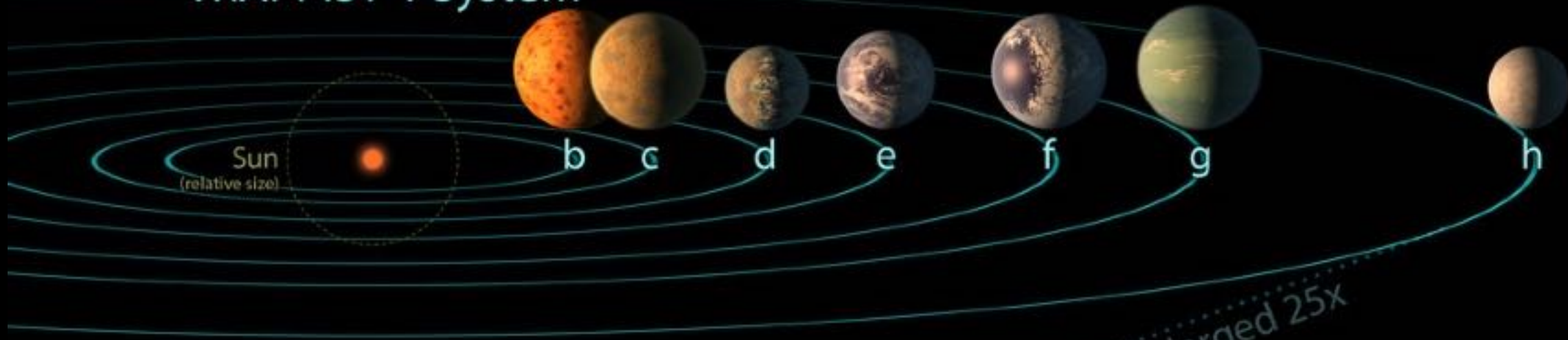
O

“The M-dwarf Opportunity”

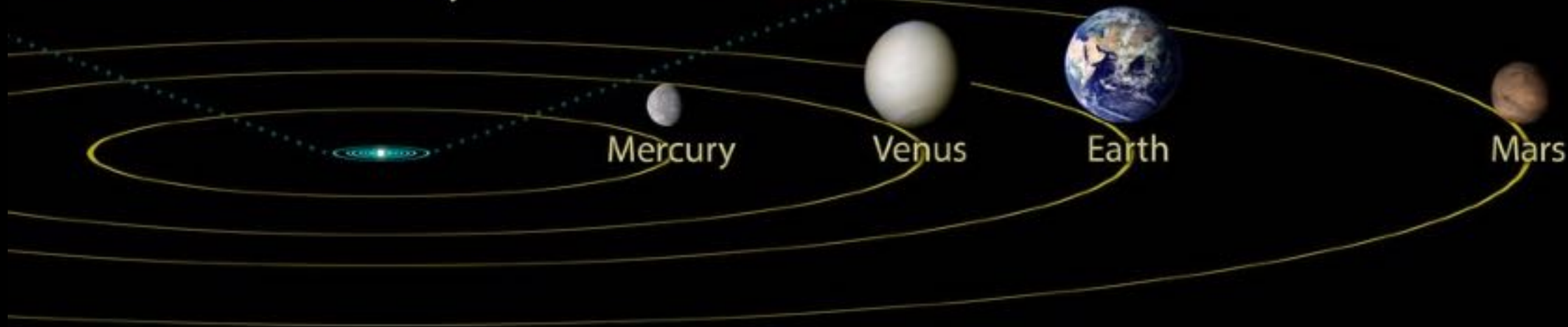




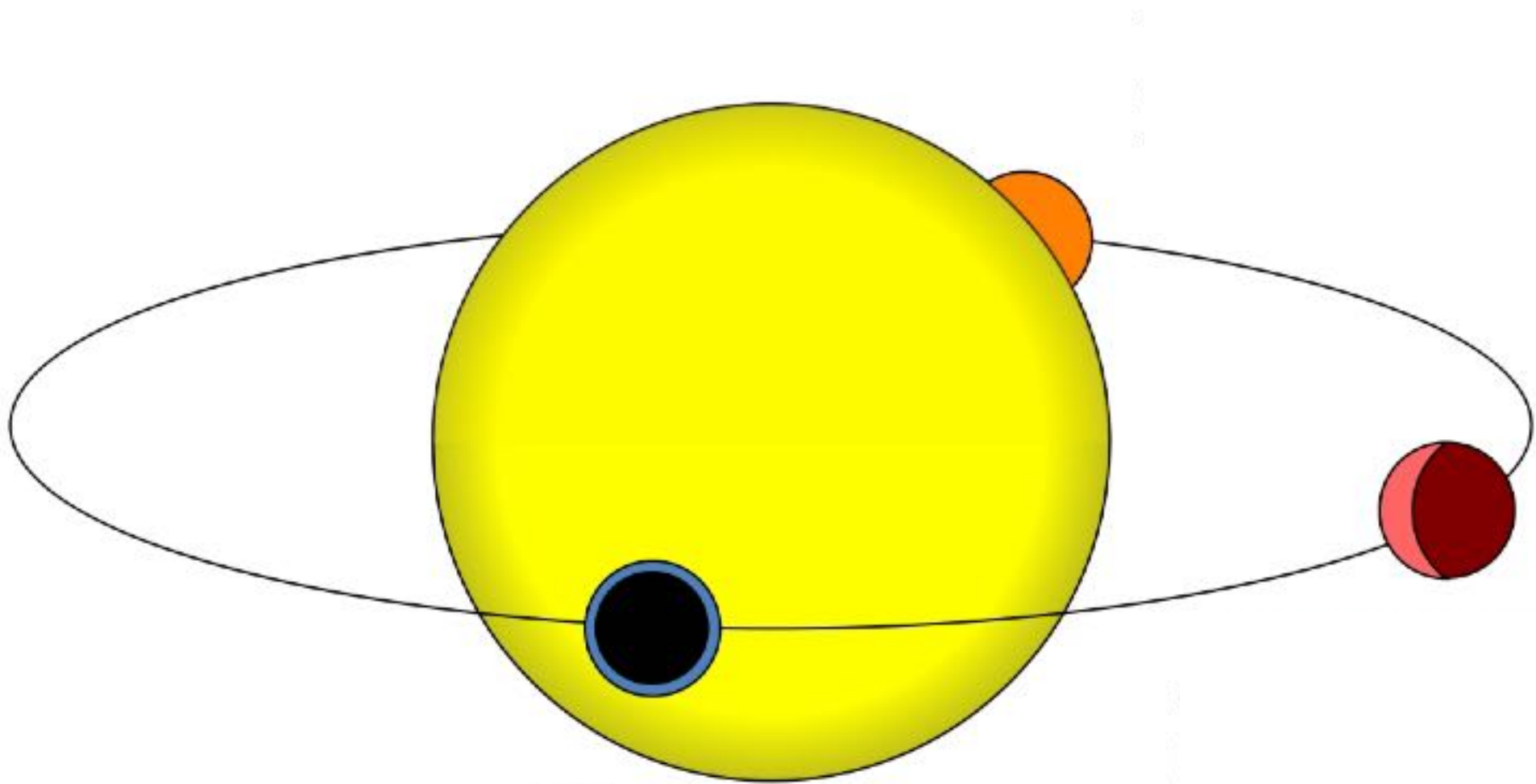
TRAPPIST-1 System



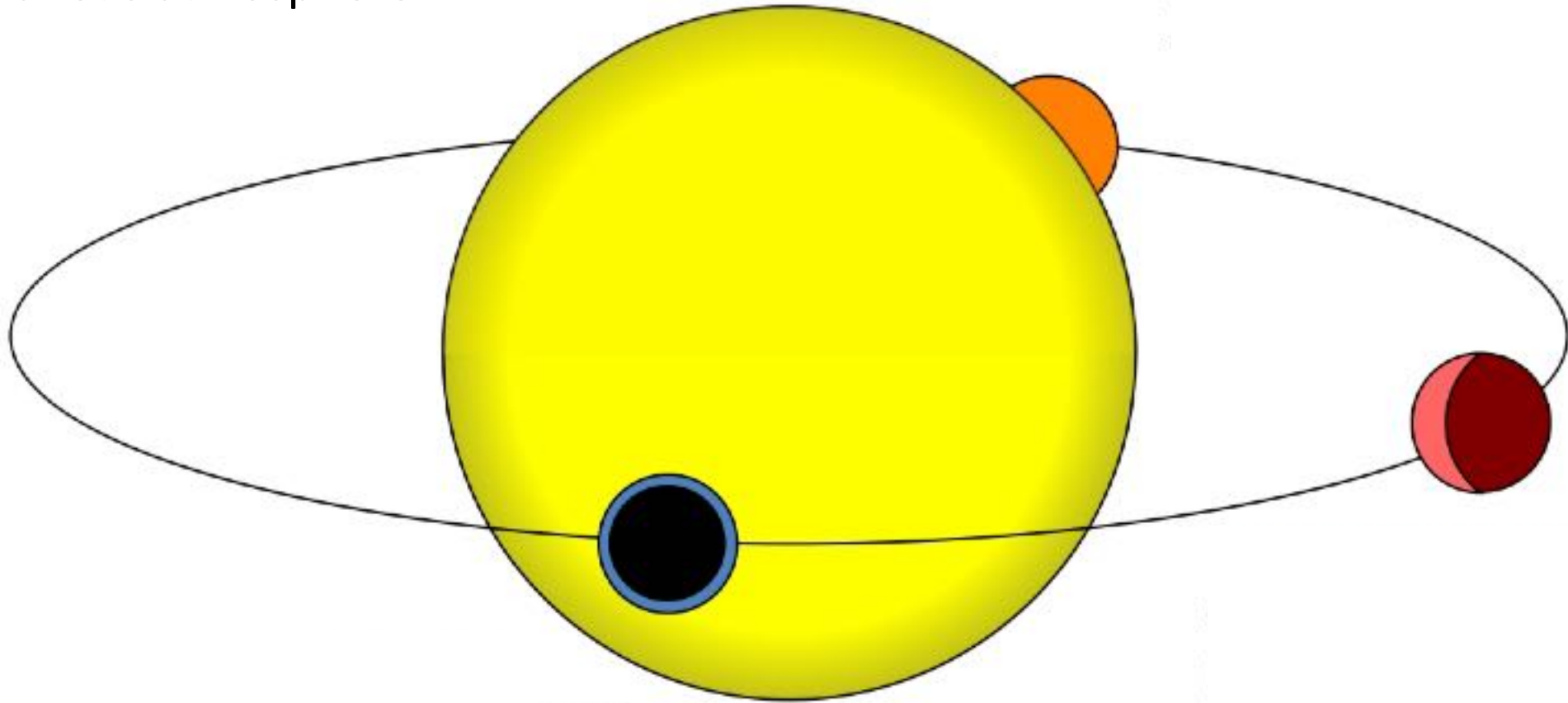
Inner Solar System



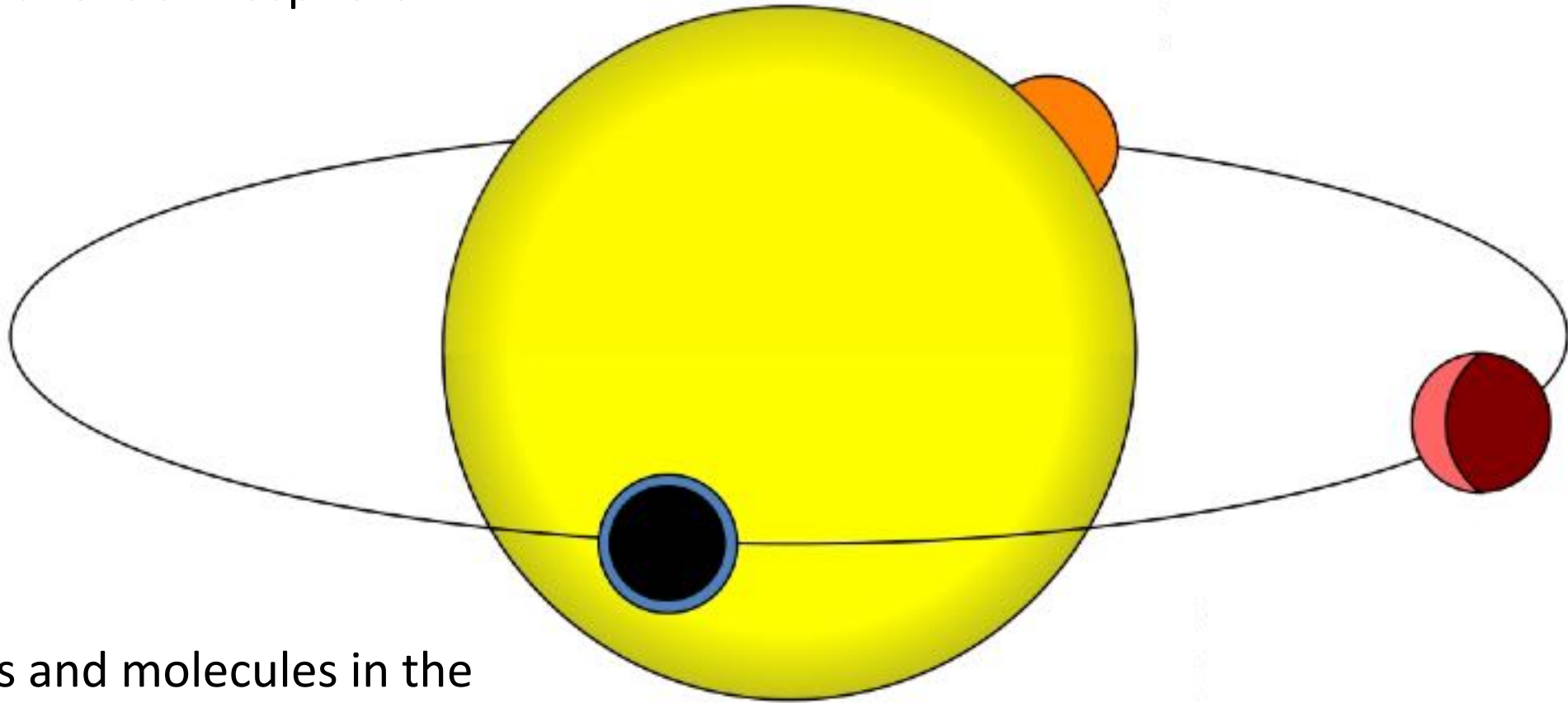
Orbits Enlarged 25x



Starlight filters through
the planet's atmosphere.

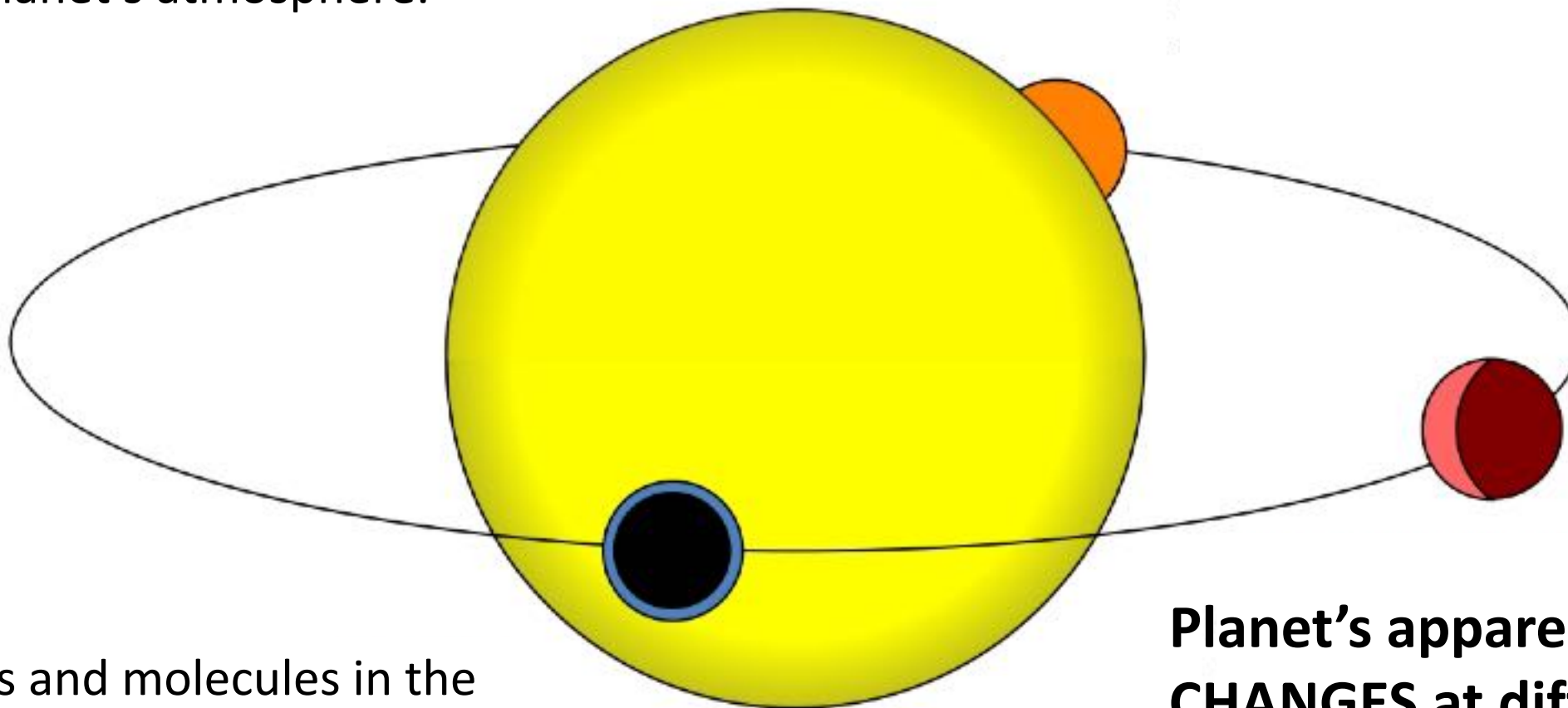


Starlight filters through the planet's atmosphere.



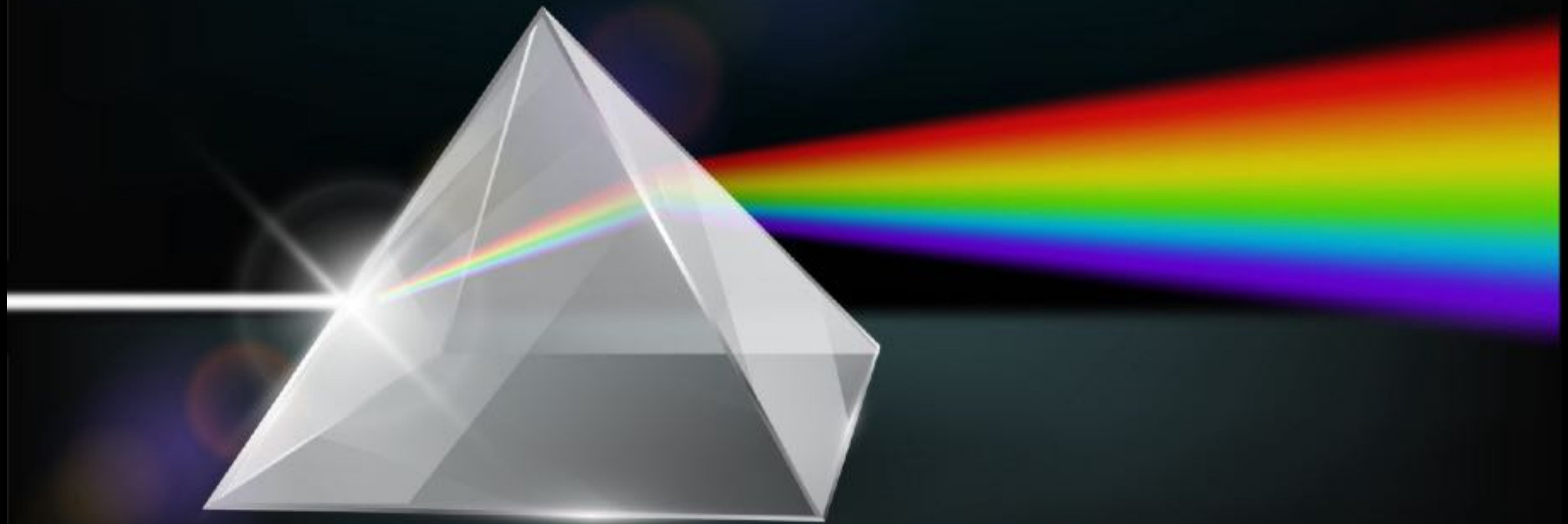
Atoms and molecules in the planet's atmosphere imprint a signature in the light we see.

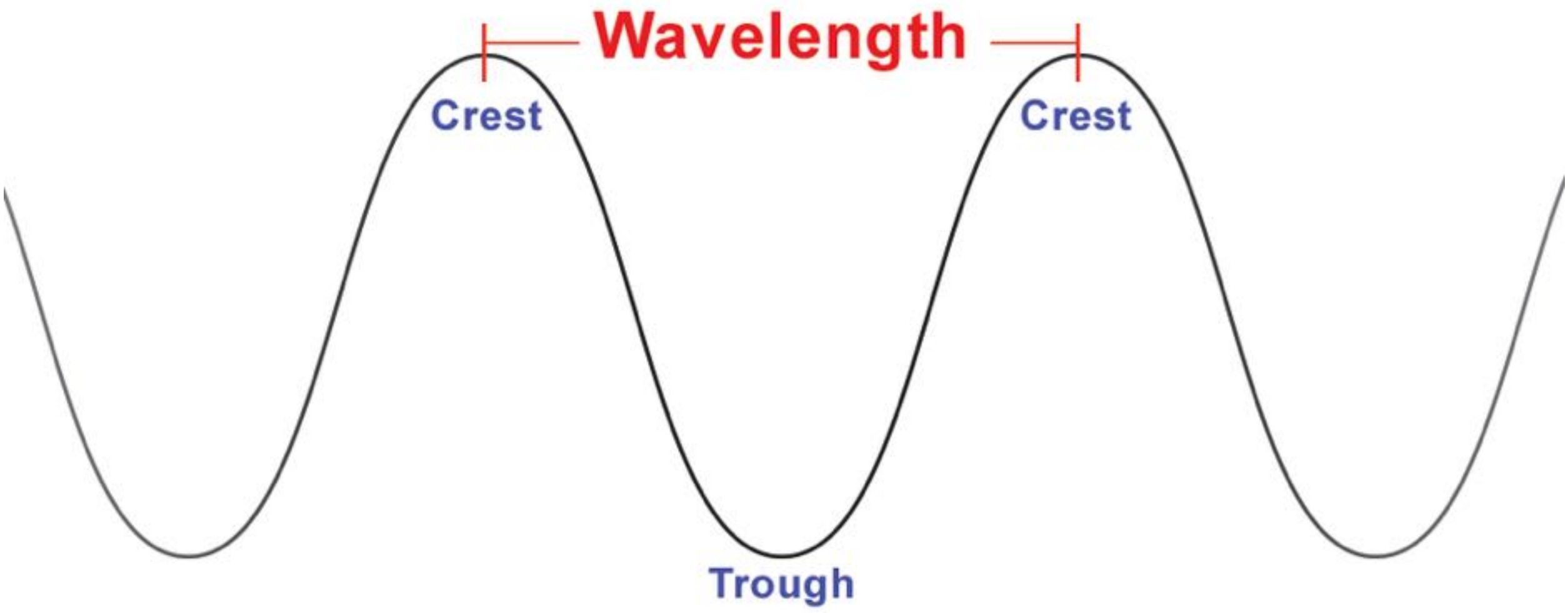
Starlight filters through the planet's atmosphere.



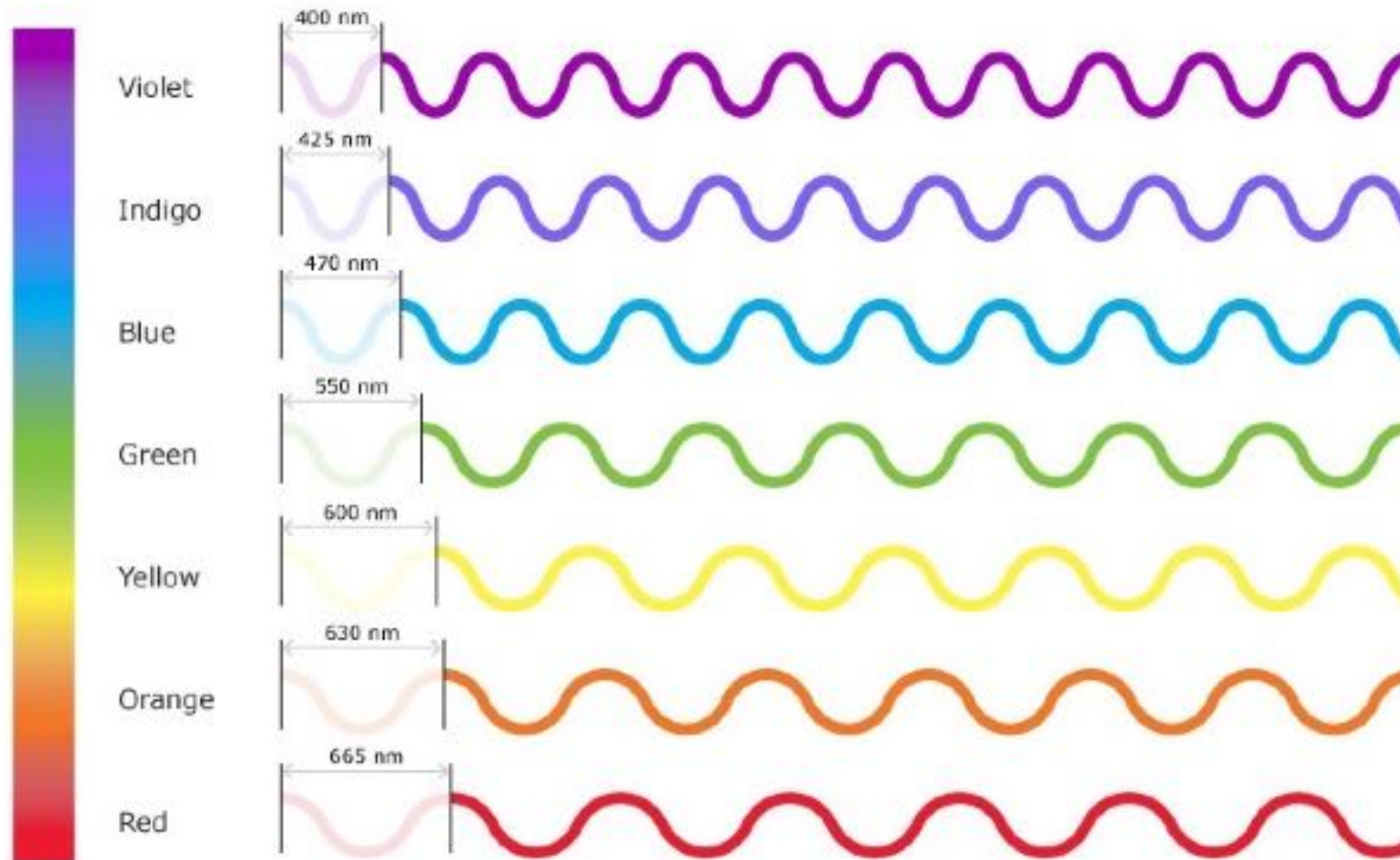
Atoms and molecules in the planet's atmosphere imprint a signature in the light we see.

**Planet's apparent radius
CHANGES at different
wavelengths**

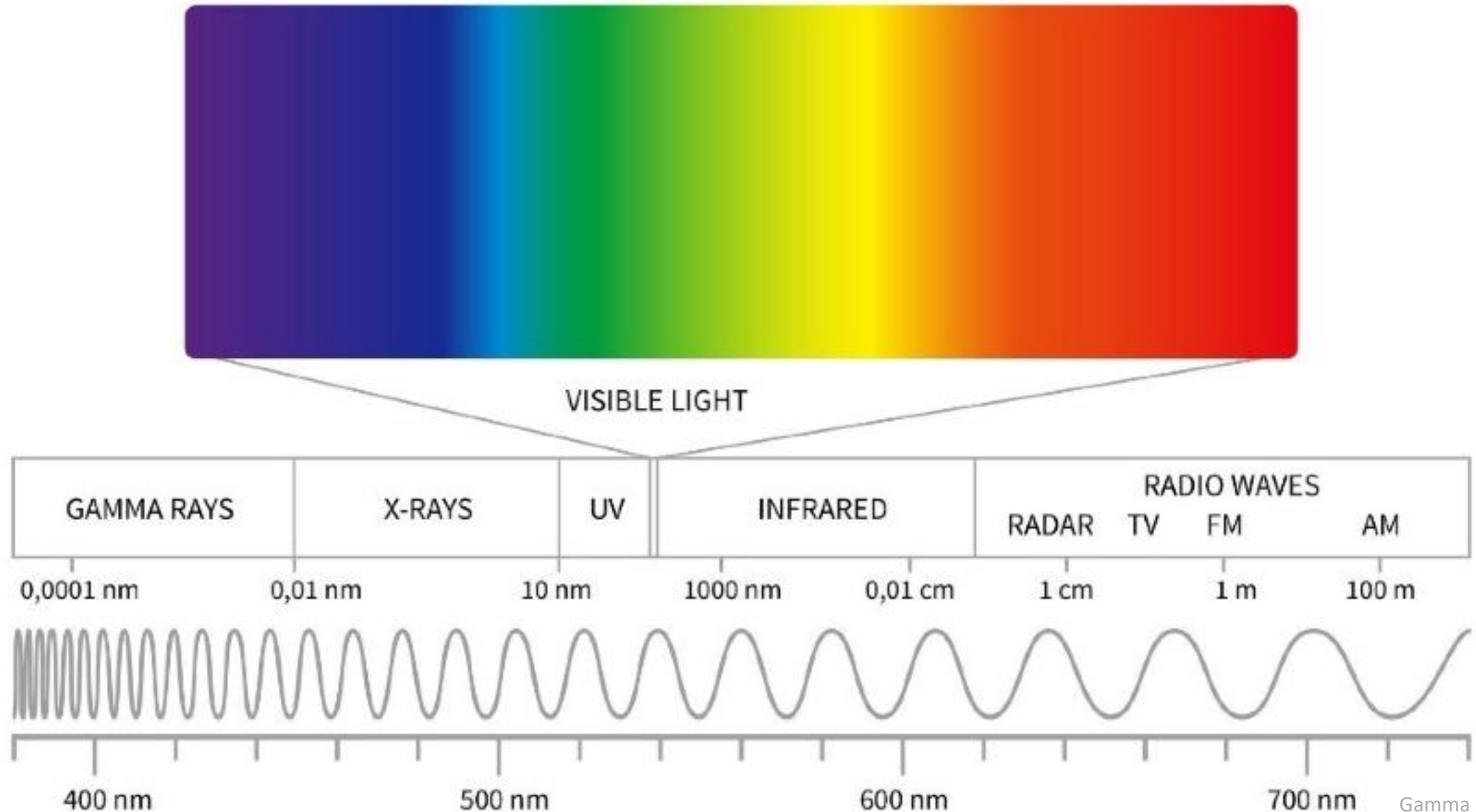




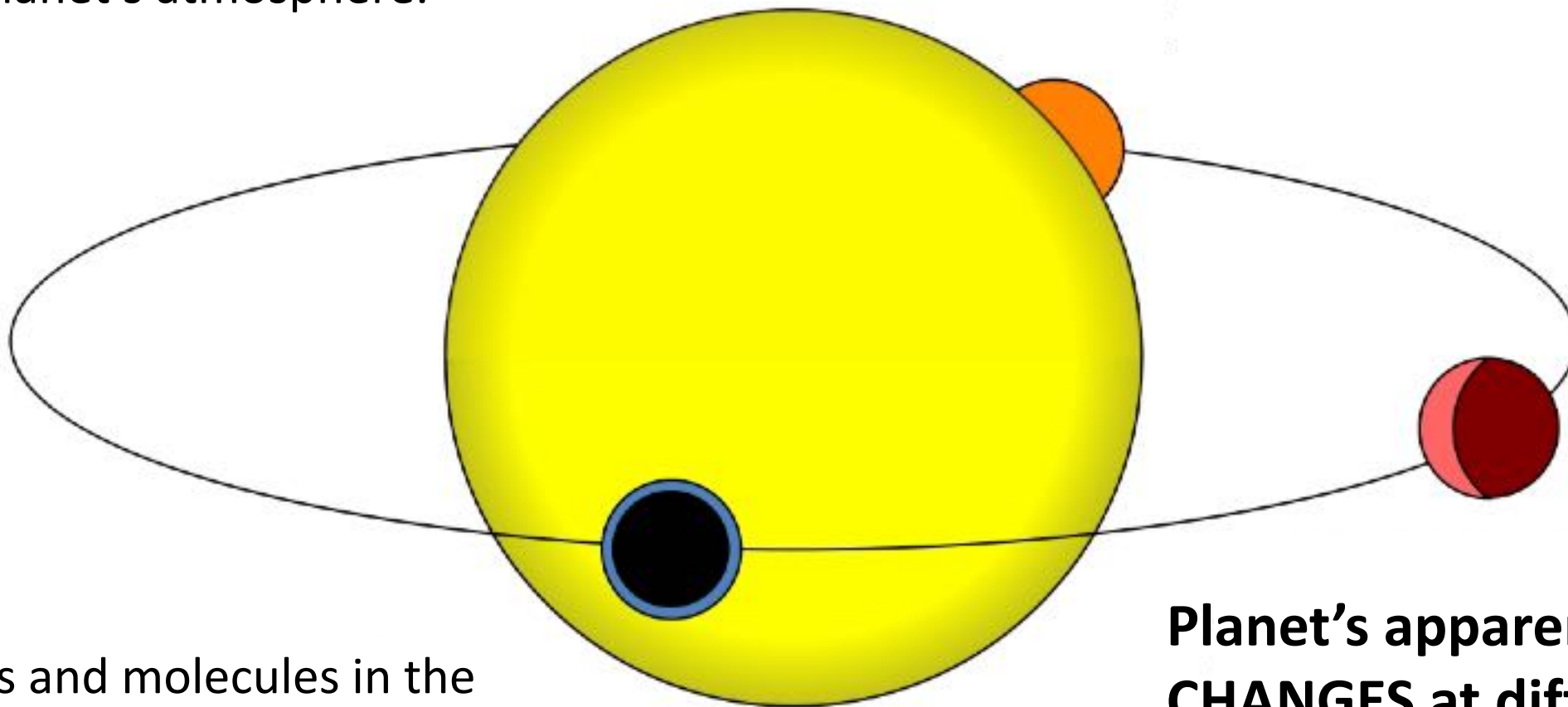
Light is a wave, with different colors representing different wavelengths



This applies across the entire **electromagnetic spectrum**

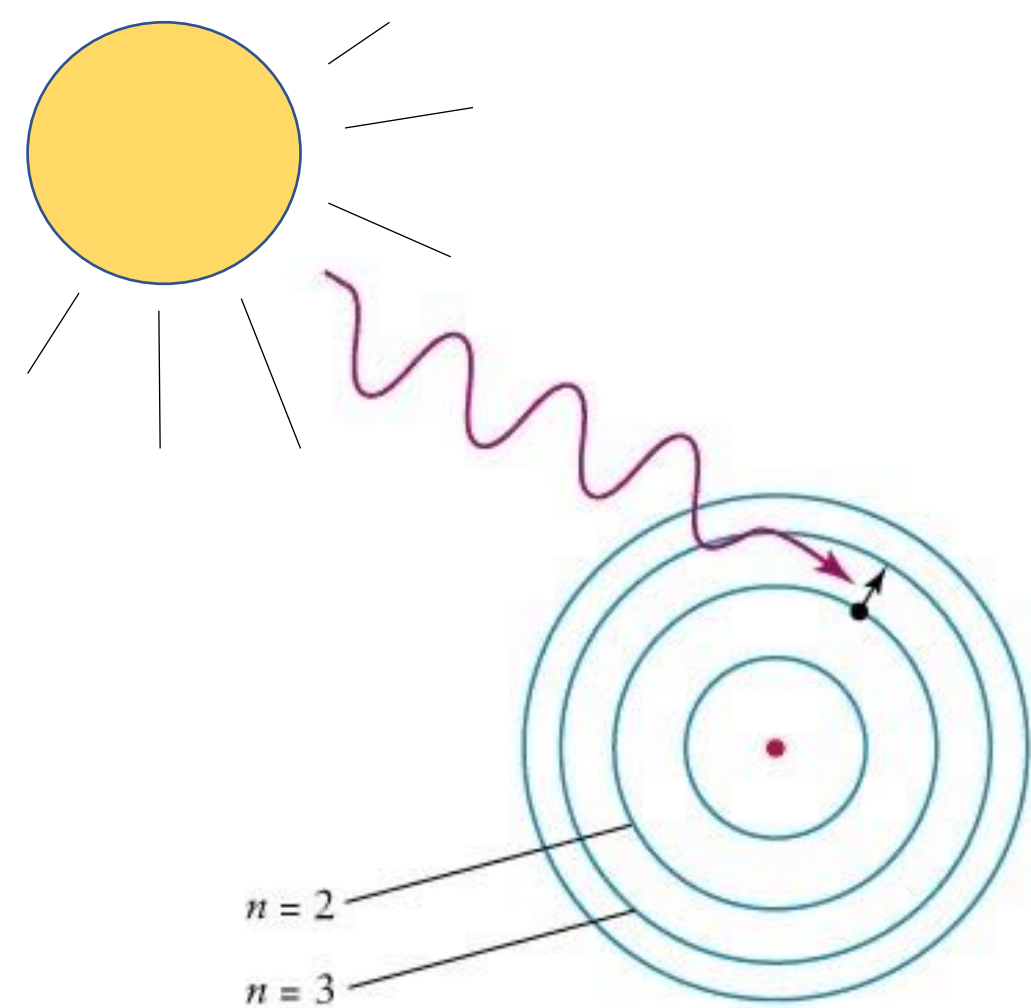


Starlight filters through the planet's atmosphere.

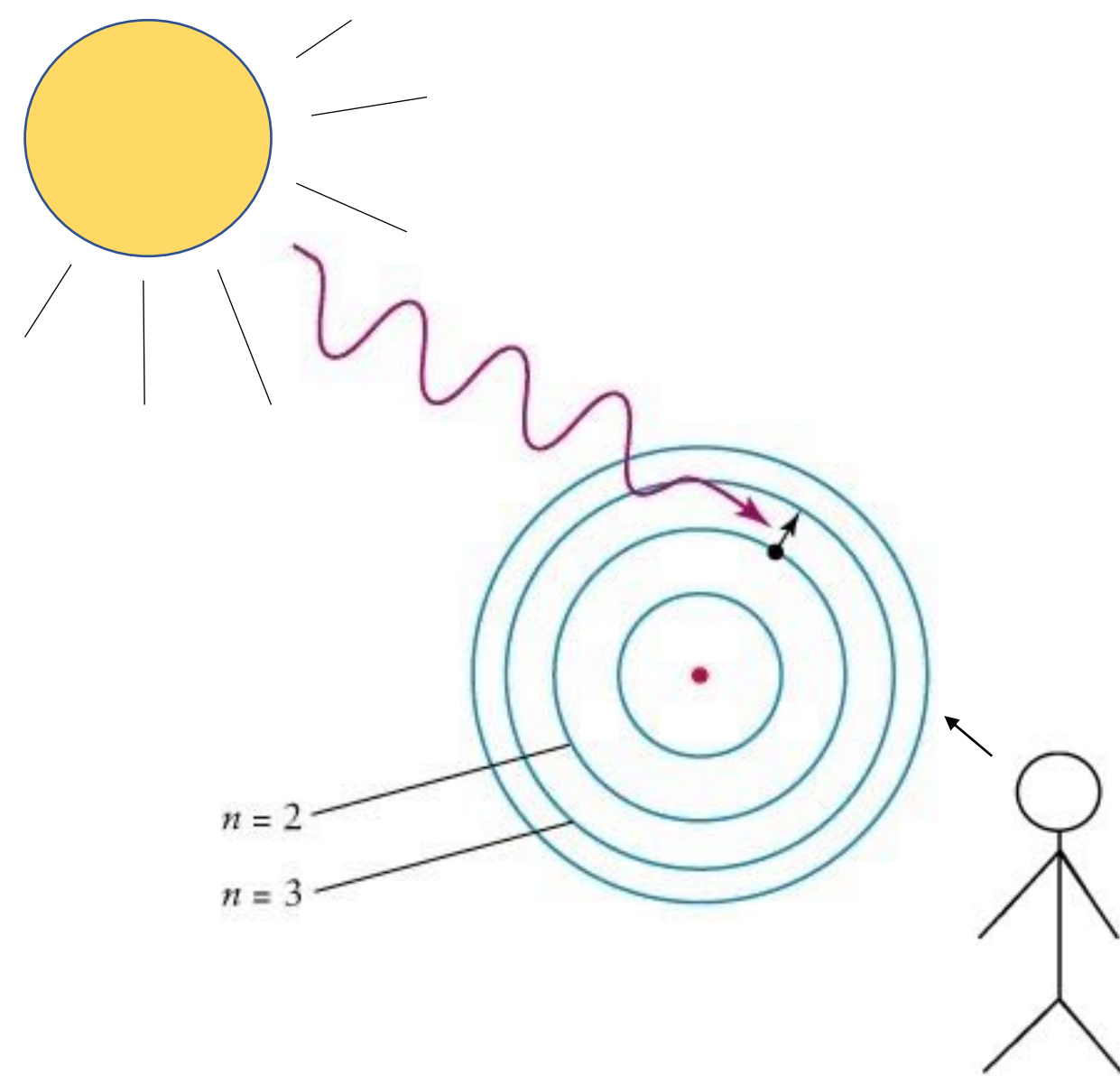


Atoms and molecules in the planet's atmosphere imprint a signature in the light we see.

**Planet's apparent radius
CHANGES at different
wavelengths**



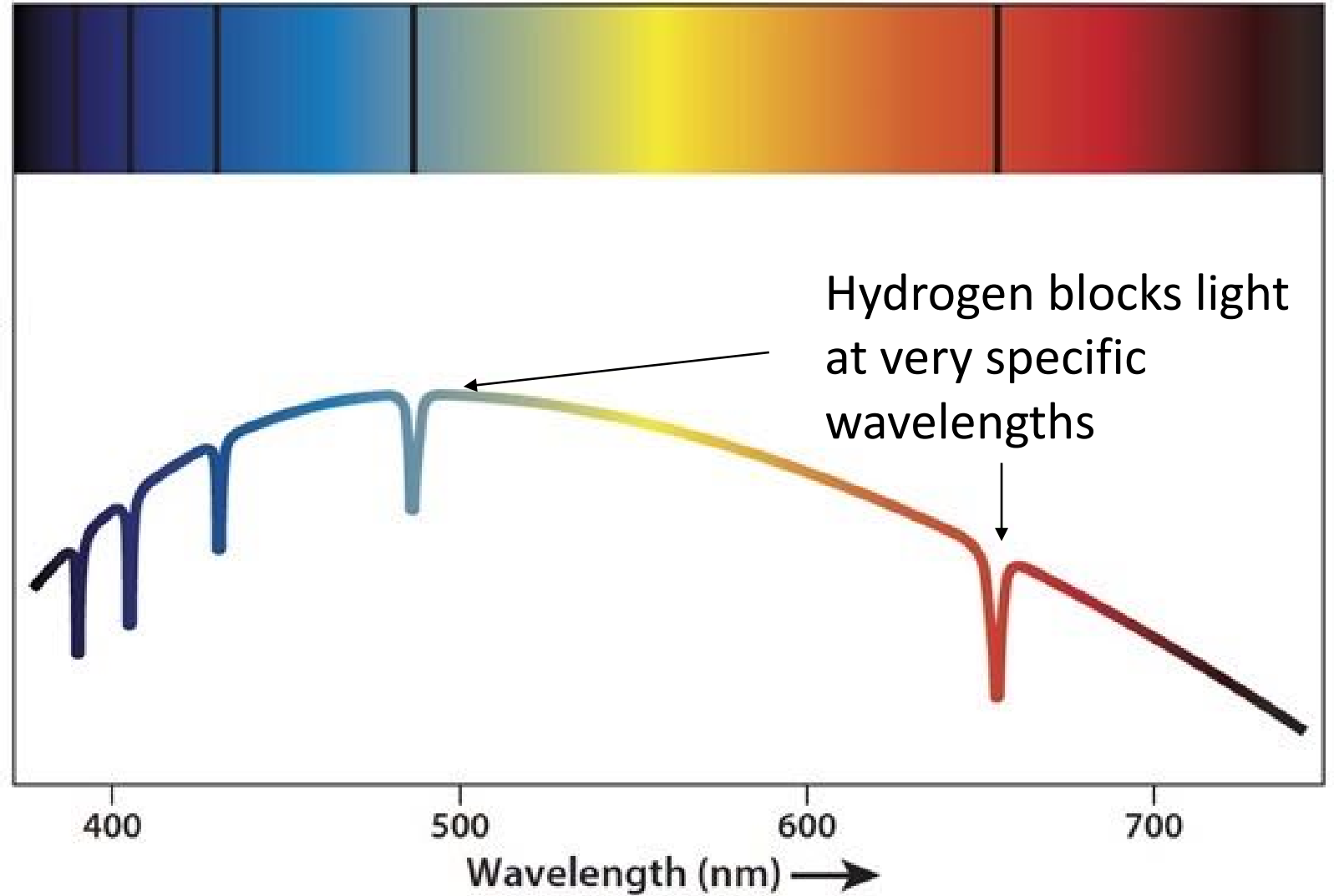
Atoms and molecules absorb light at very specific wavelengths, leaving a “chemical signature”



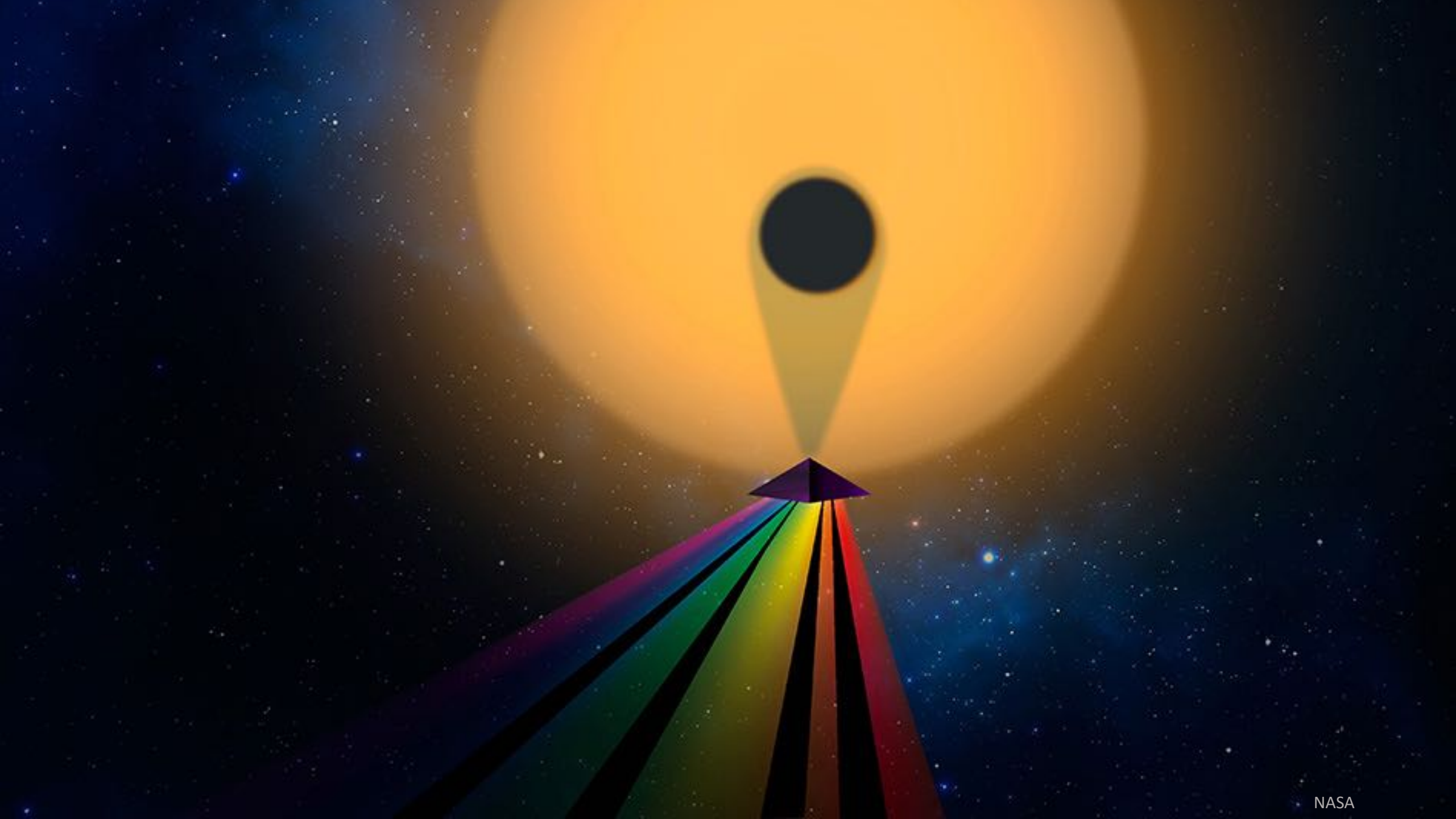
Atoms and molecules absorb light at very specific wavelengths, leaving a “chemical signature”

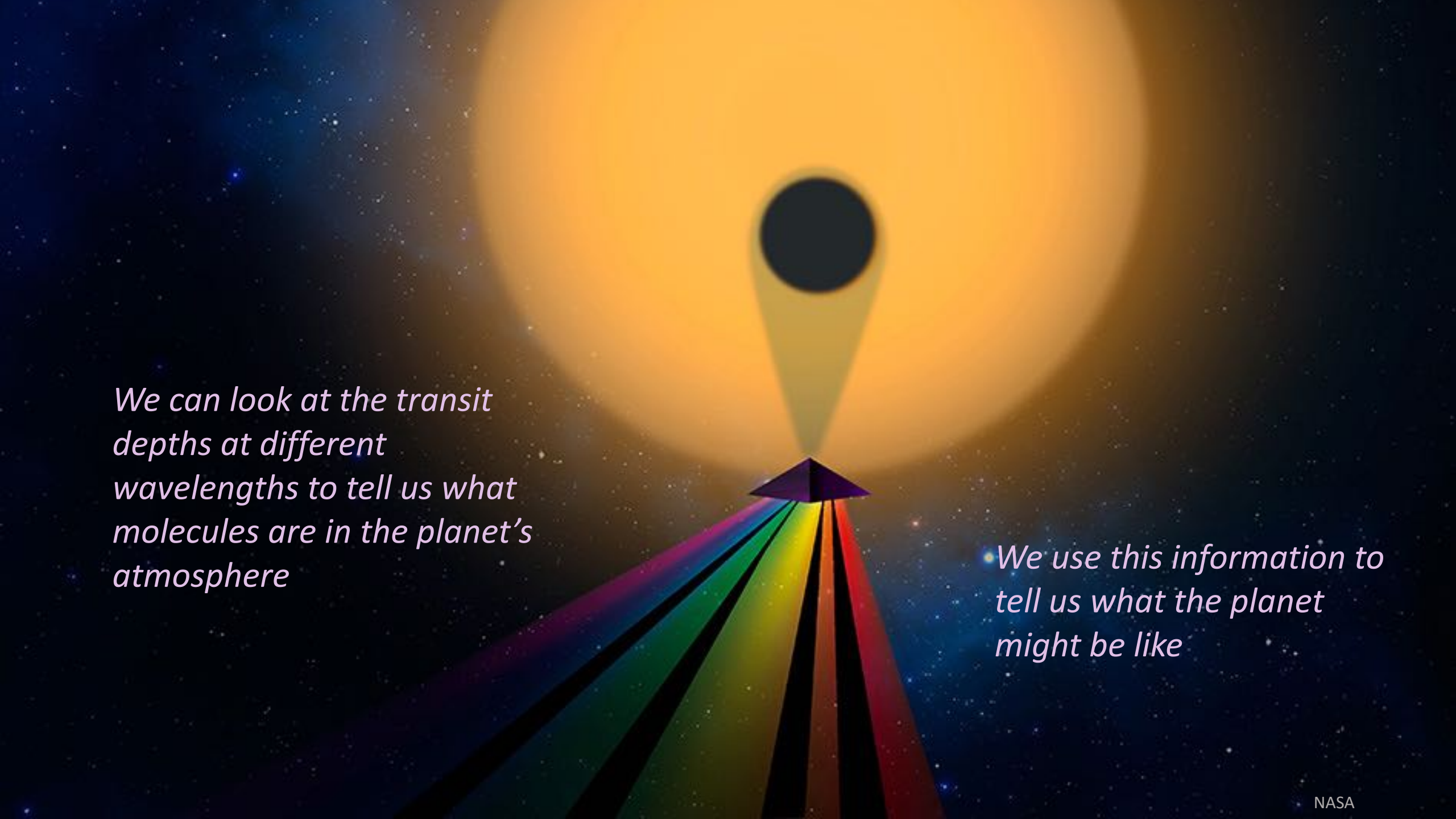
An observer sees light being blocked at very specific wavelengths

Hydrogen absorption spectrum



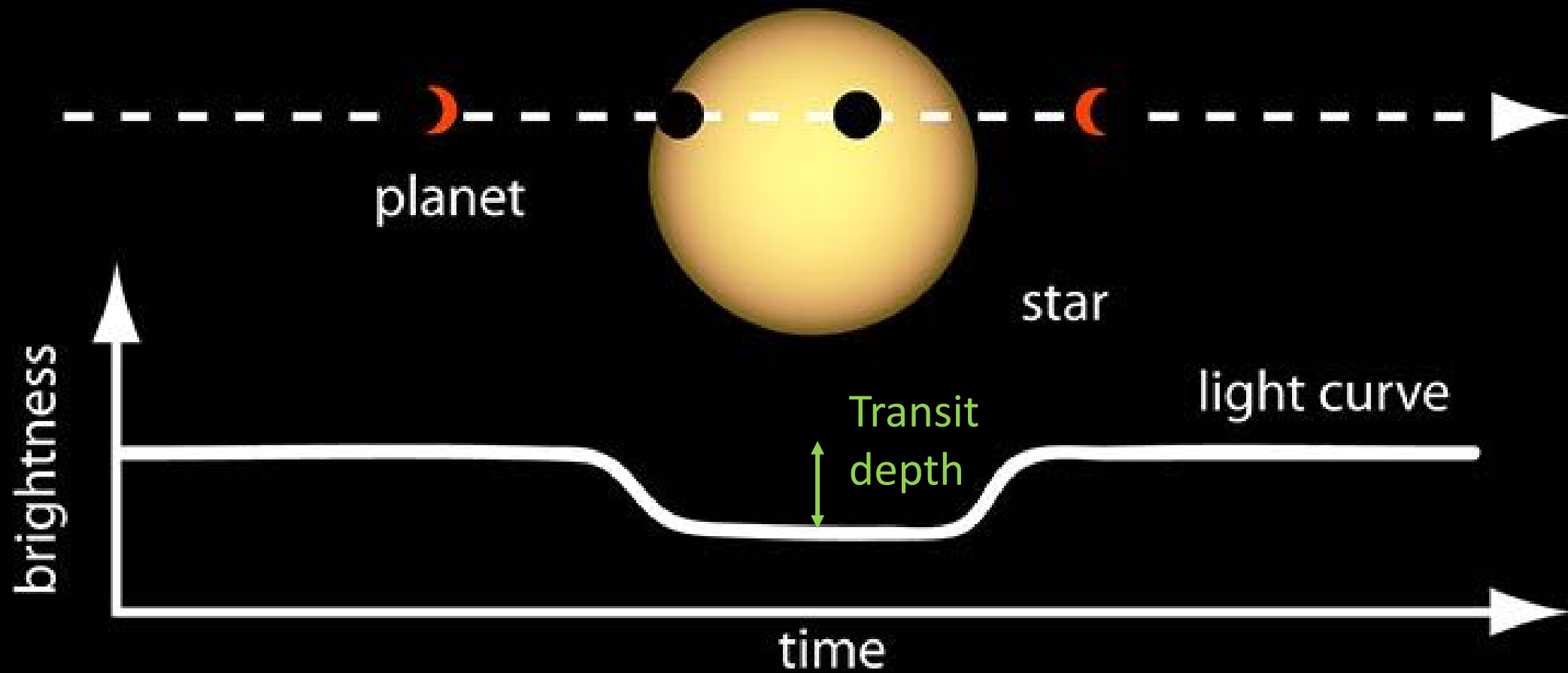
Spectrum: the measurement of the amount of light at different wavelengths



A diagram illustrating the process of transit spectroscopy. A large, bright yellow sun is at the top center. A smaller, dark planet is shown in transit across the sun's face. A telescope, represented by a purple pyramid, is positioned below the sun. A beam of light from the sun passes through the planet and the telescope. This light is then dispersed into a spectrum of colors, shown as a fan of rays ranging from purple on the left to red on the right. The background is a dark blue space filled with stars.

We can look at the transit depths at different wavelengths to tell us what molecules are in the planet's atmosphere

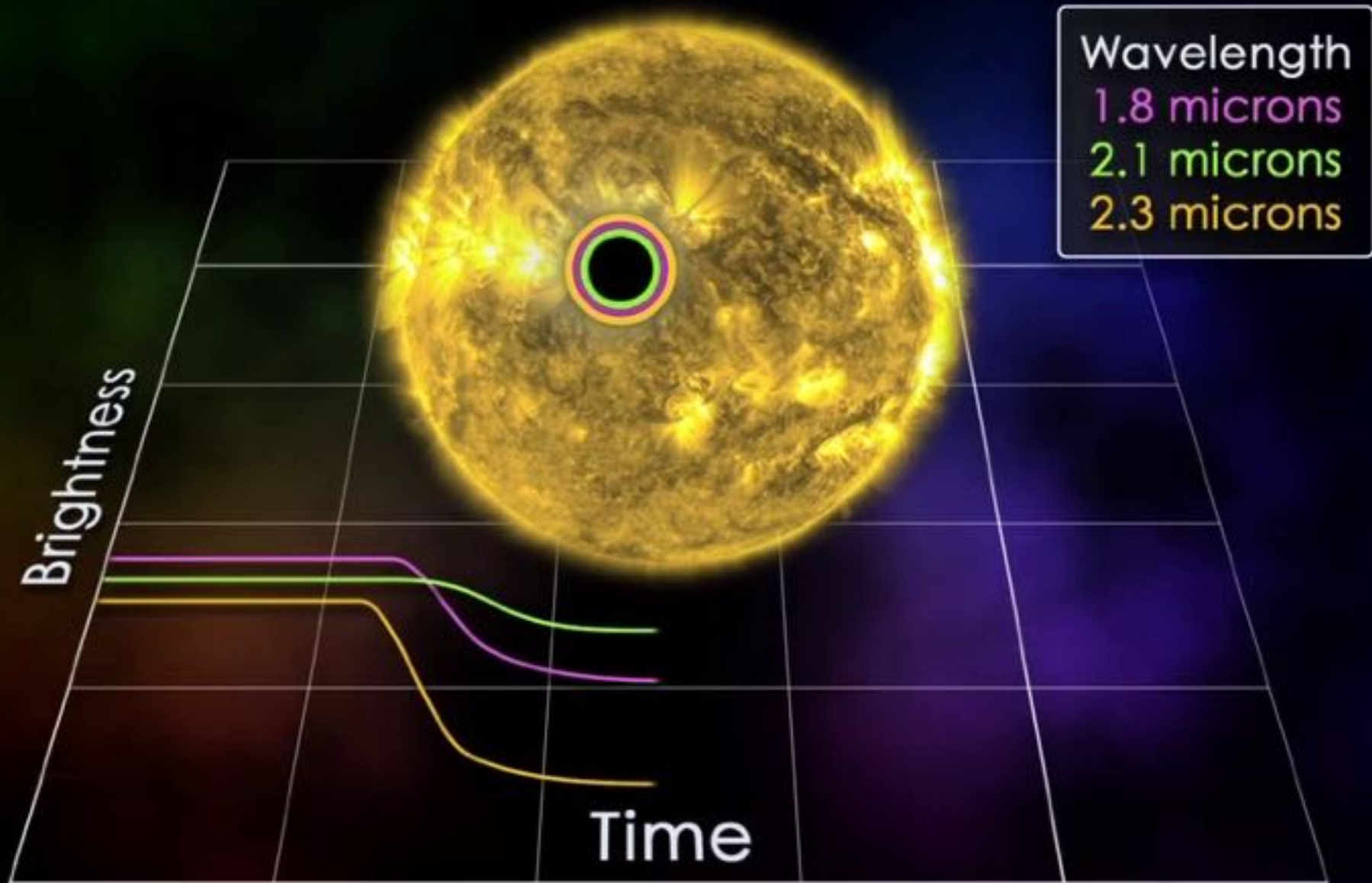
We use this information to tell us what the planet might be like



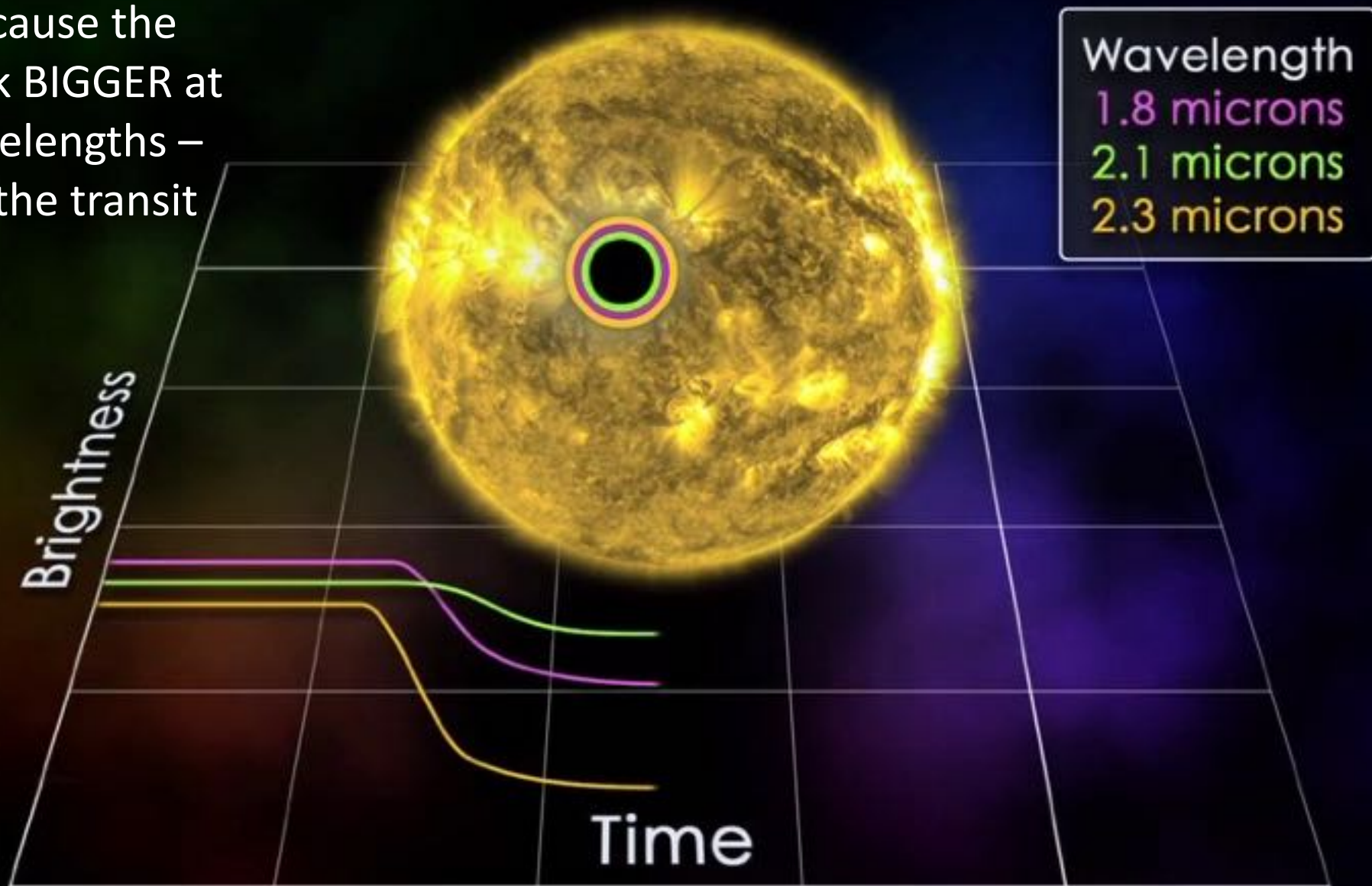
Transmission Spectroscopy

Transit depth is actually a function of wavelength

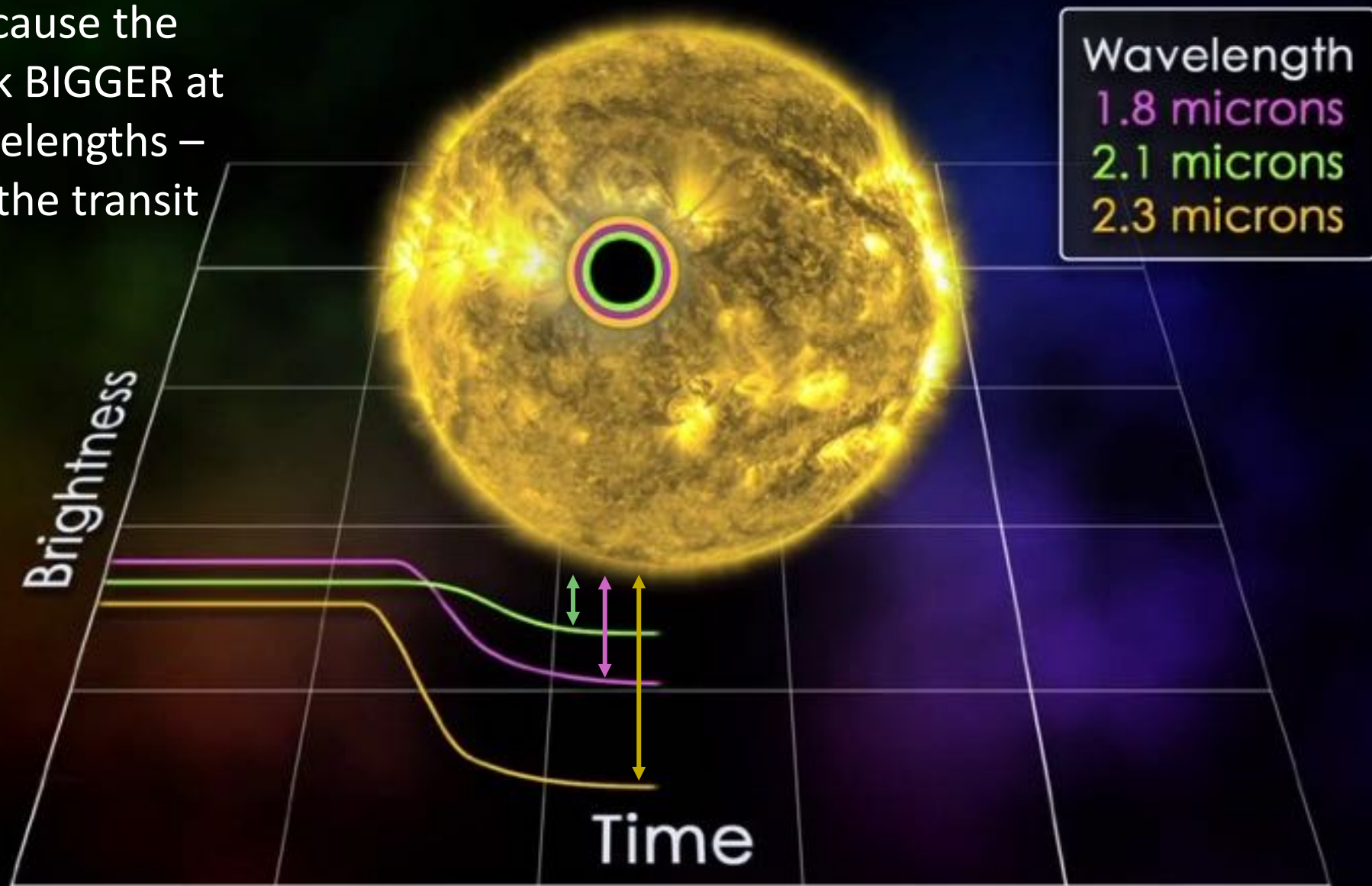




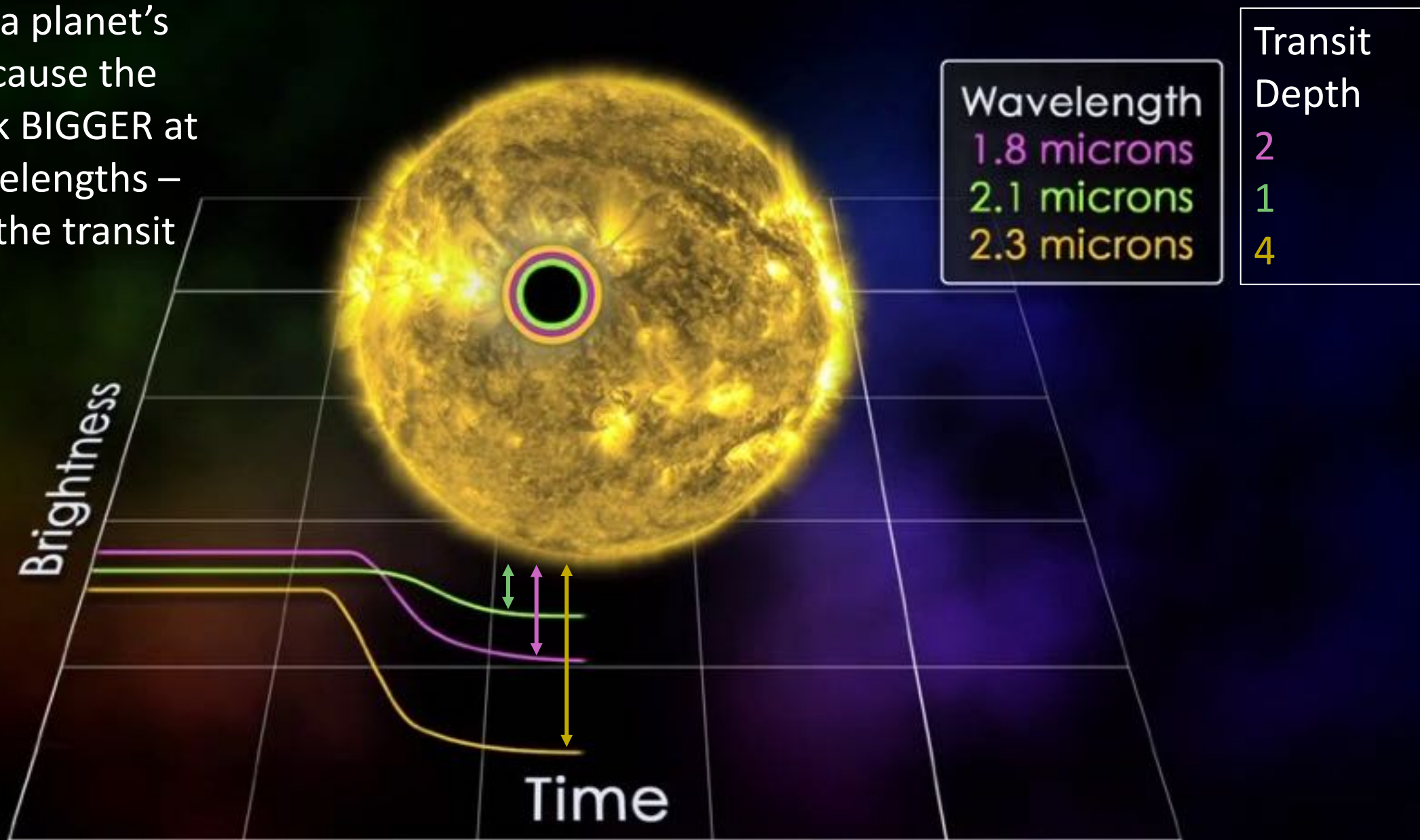
Molecules in a planet's atmosphere cause the planet to look BIGGER at different wavelengths – this changes the transit depth



Molecules in a planet's atmosphere cause the planet to look BIGGER at different wavelengths – this changes the transit depth

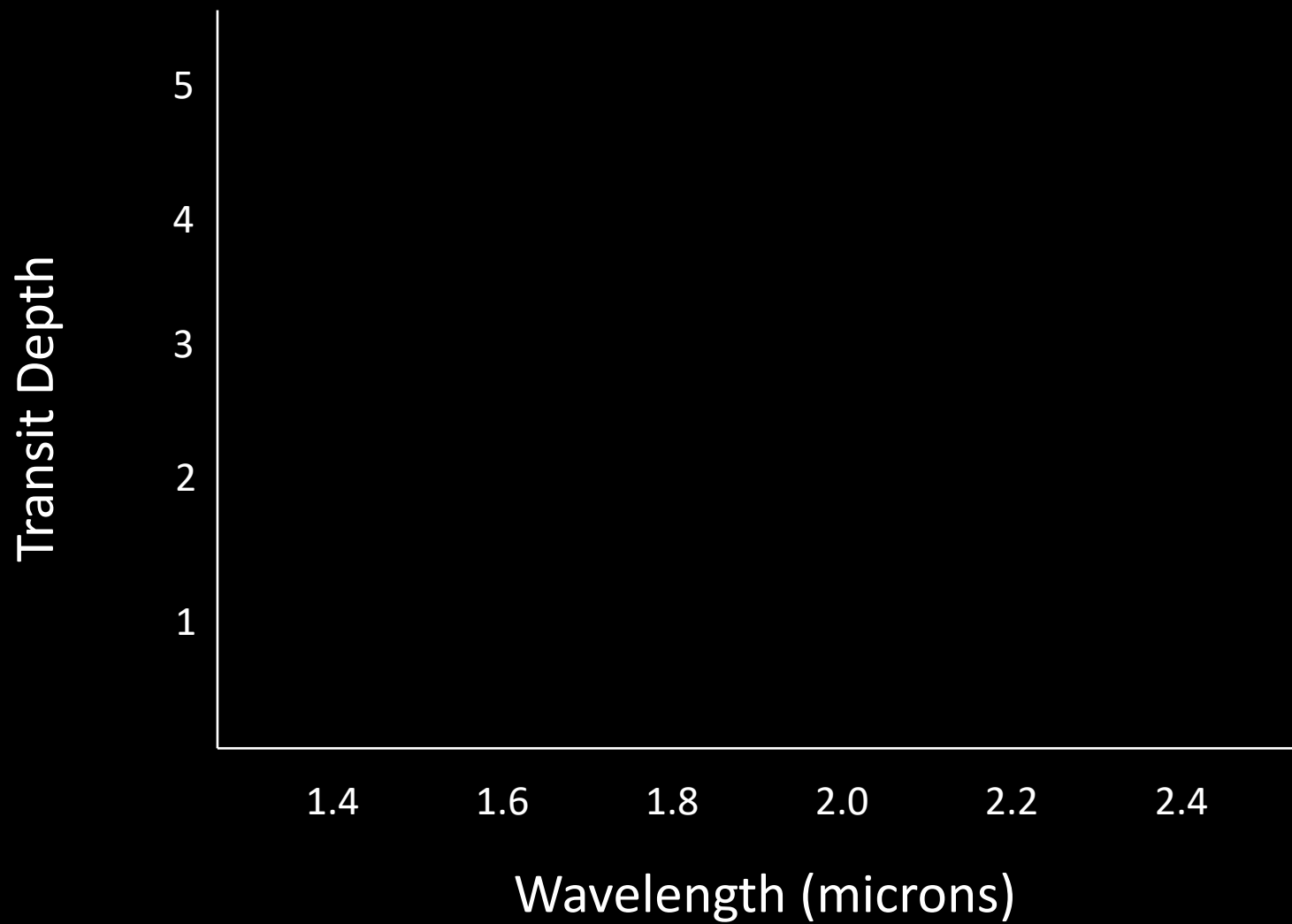


Molecules in a planet's atmosphere cause the planet to look BIGGER at different wavelengths – this changes the transit depth



Wavelength
1.8 microns
2.1 microns
2.3 microns

Transit
Depth
2
1
4



Wavelength

1.8 microns

2.1 microns

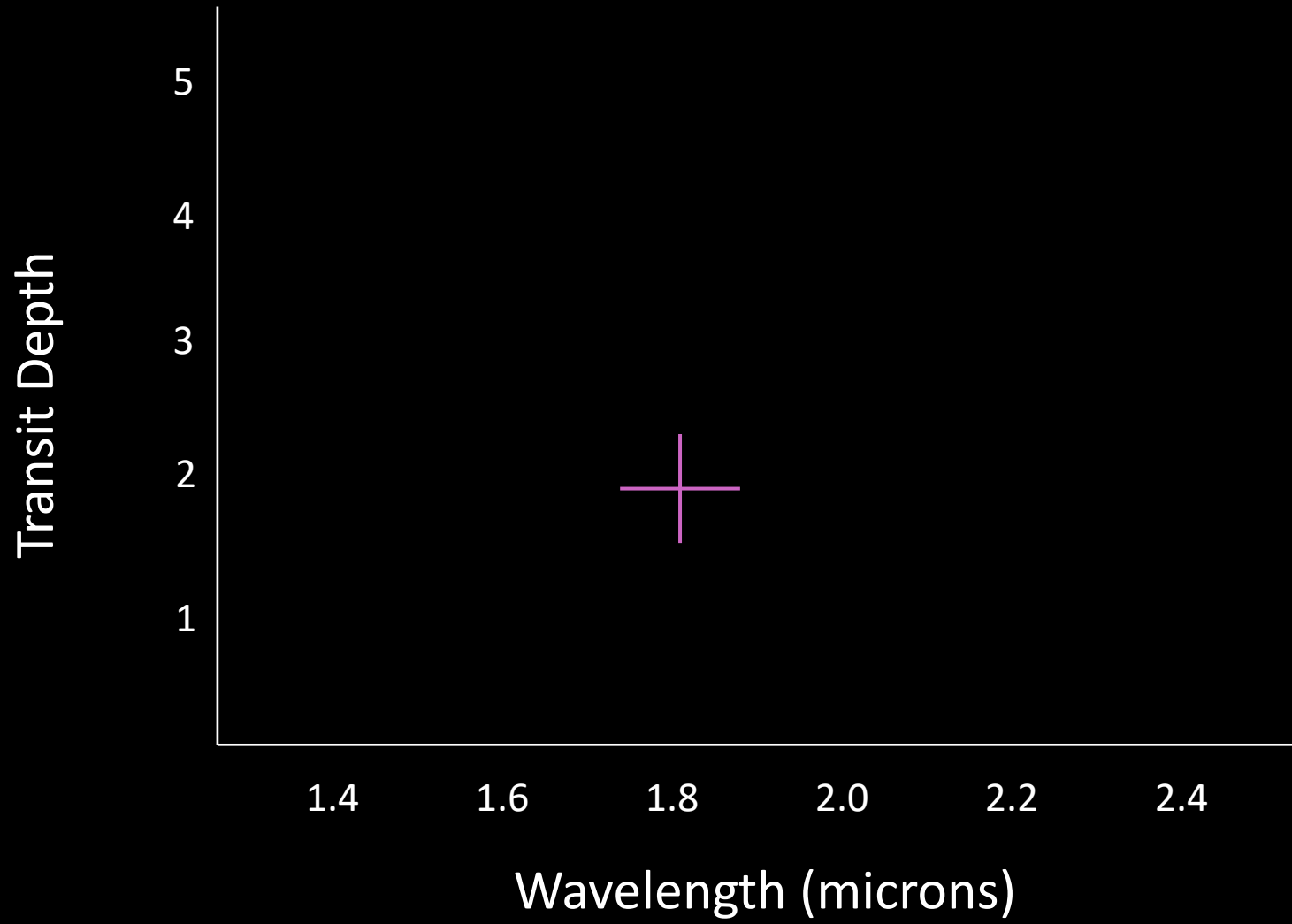
2.3 microns

Transit
Depth

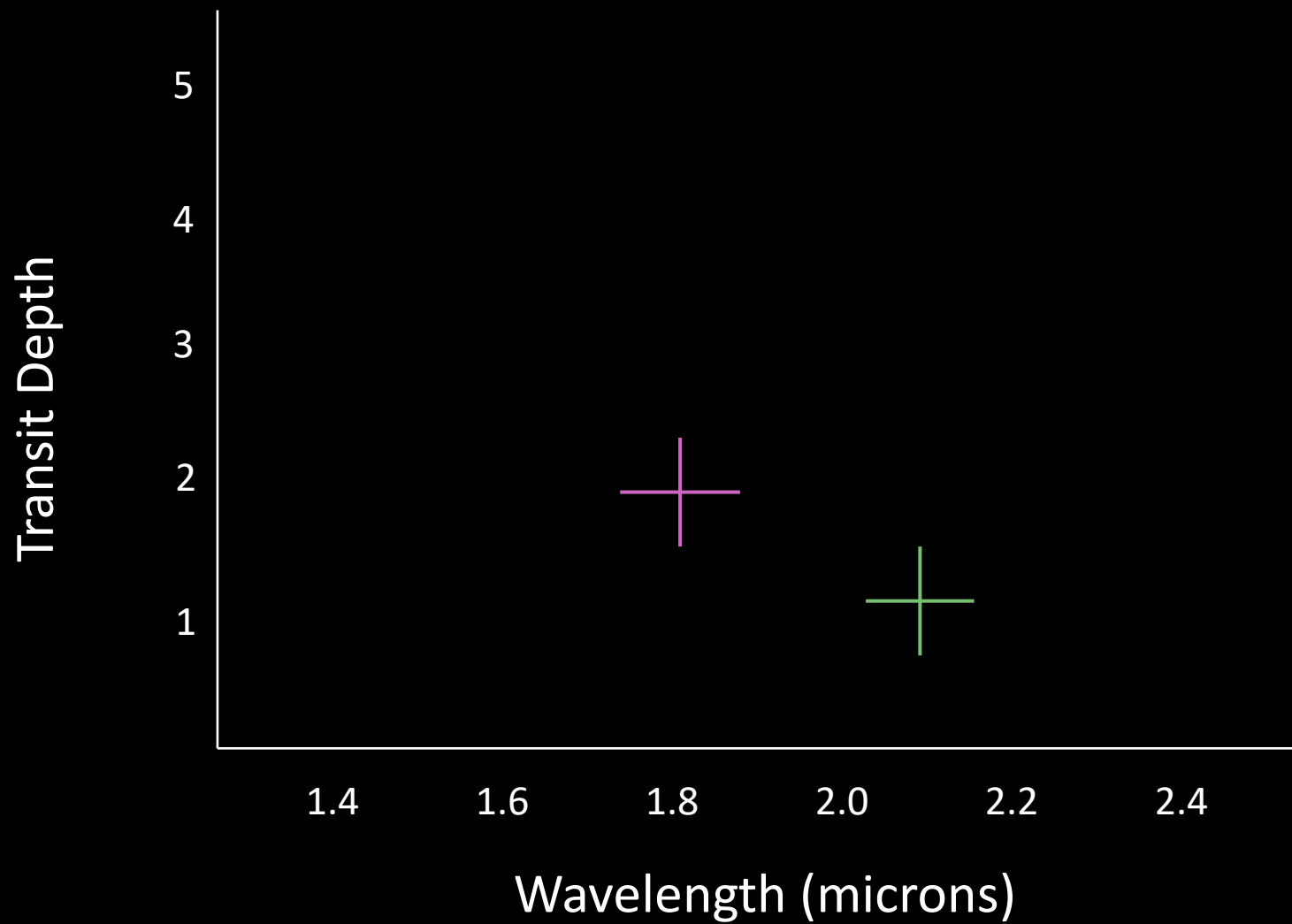
2

1

4

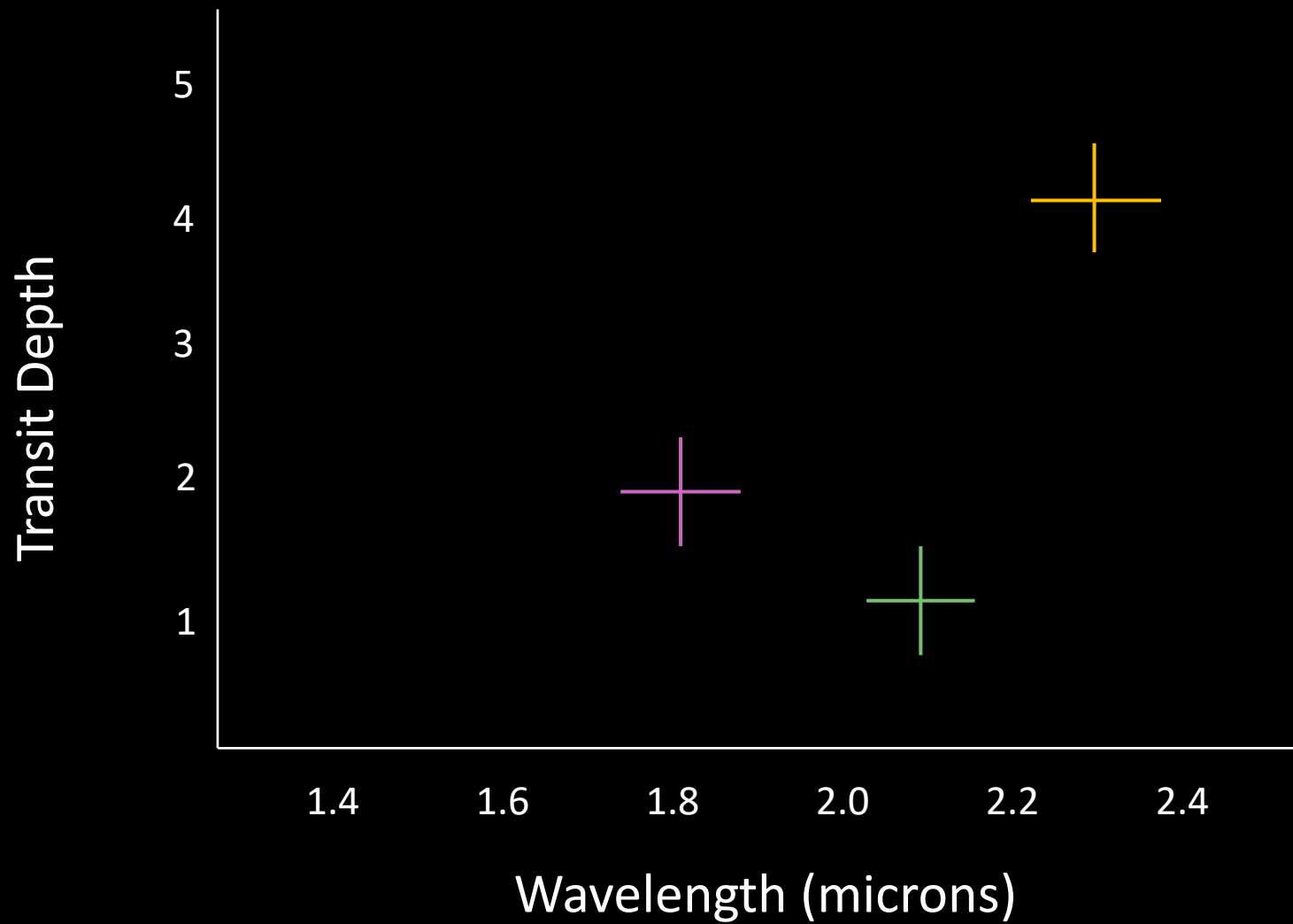


Wavelength	Transit Depth
1.8 microns	2
2.1 microns	1
2.3 microns	4



Wavelength
1.8 microns
2.1 microns
2.3 microns

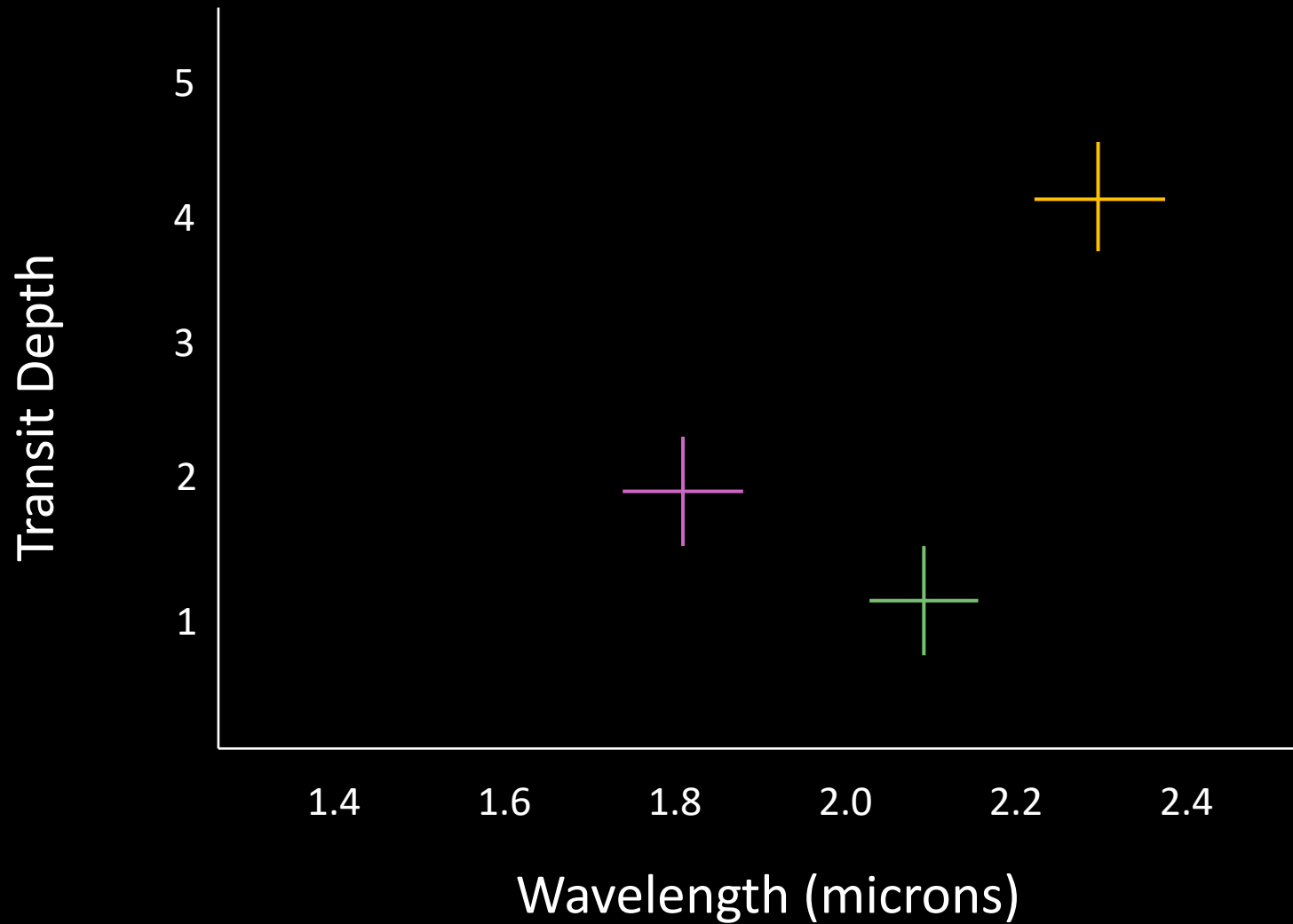
Transit
Depth
2
1
4



Wavelength
1.8 microns
2.1 microns
2.3 microns

Transit
Depth
2
1
4

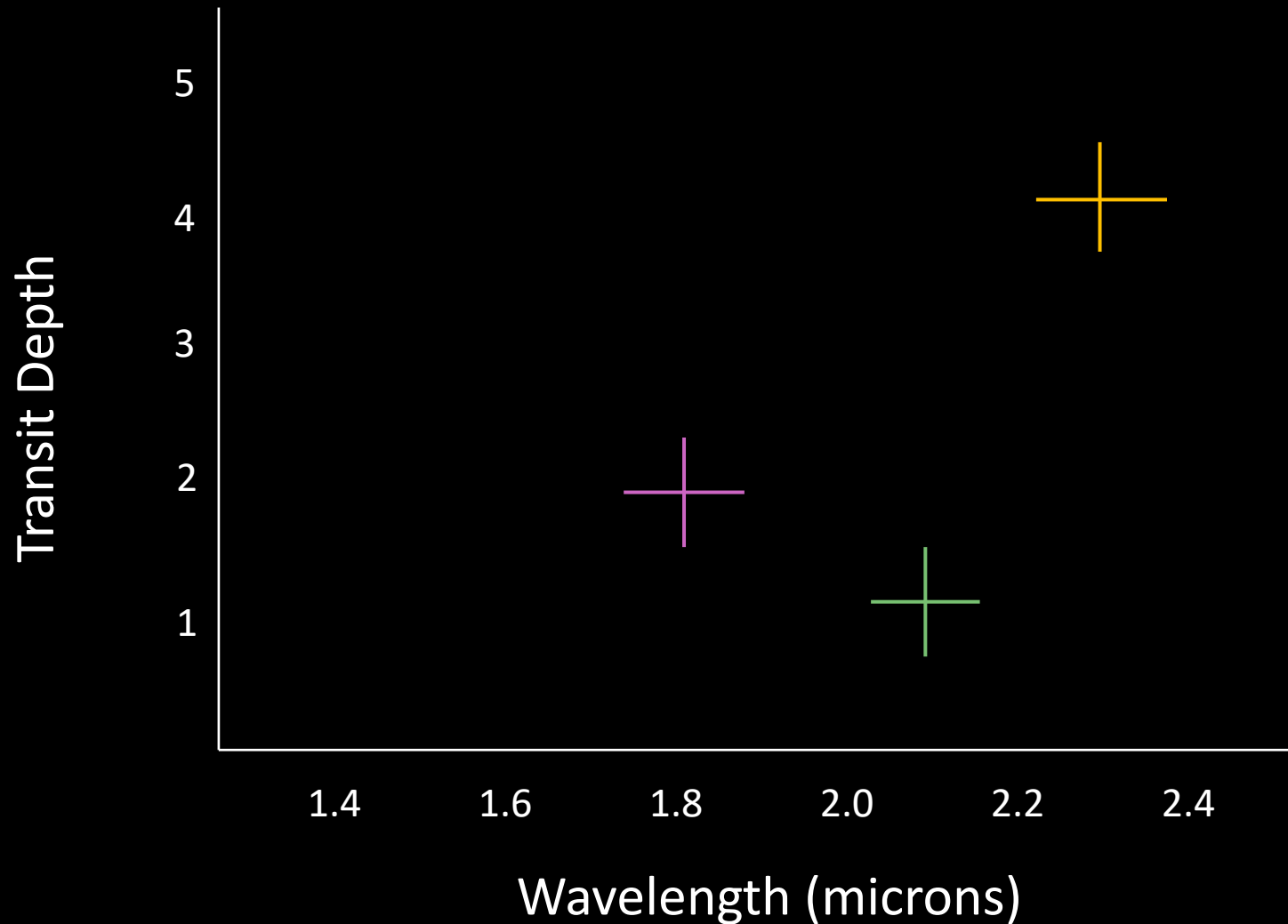
Transmission Spectrum



Wavelength
1.8 microns
2.1 microns
2.3 microns

Transit
Depth
2
1
4

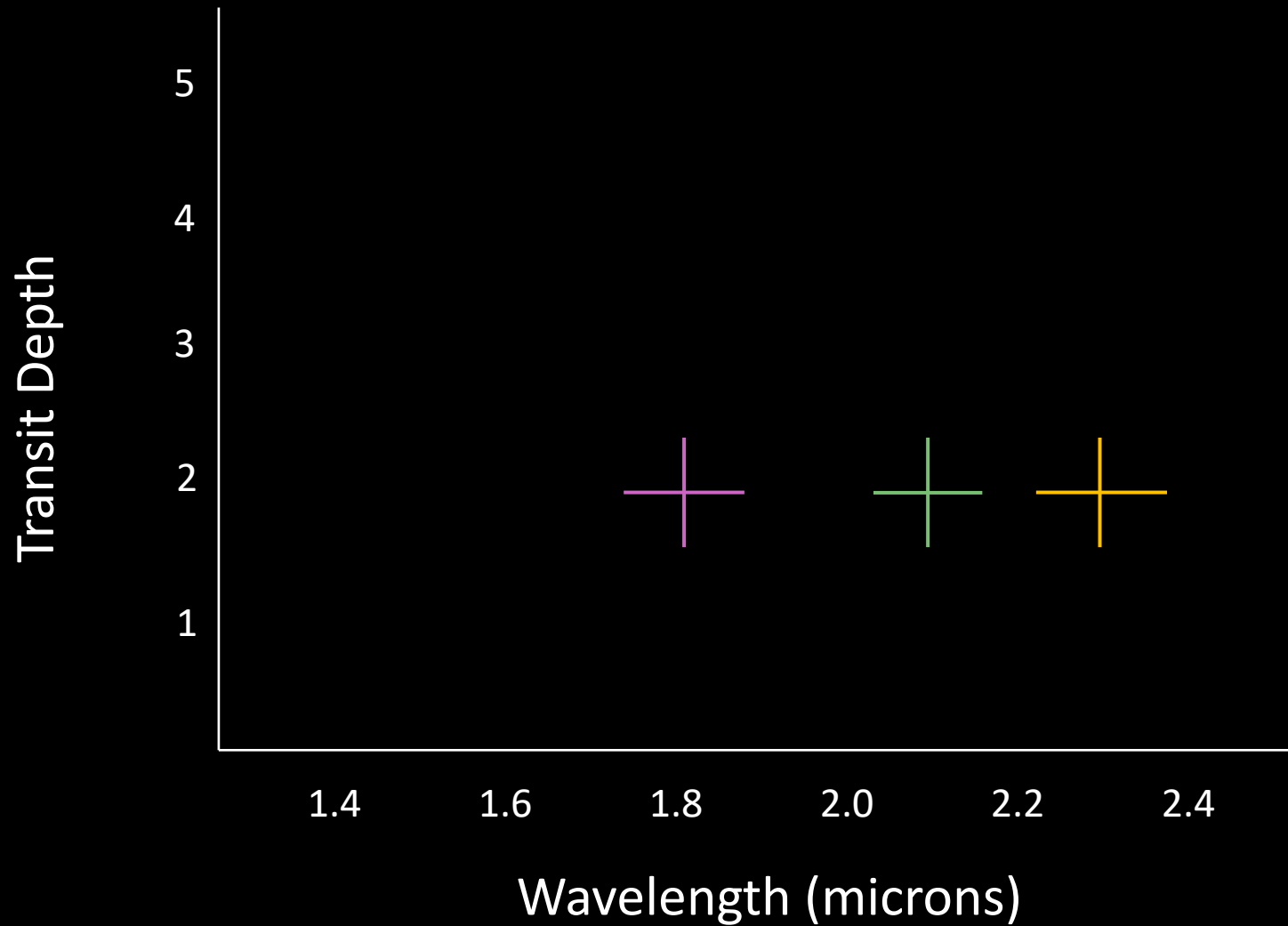
Transmission Spectrum



Wavelength	Transit Depth
1.8 microns	2
2.1 microns	1
2.3 microns	4

The fact that these points are NOT at the same transit depth means this planet HAS an atmosphere

Transmission Spectrum



Wavelength	Transit Depth
1.8 microns	2
2.1 microns	1
2.3 microns	4

If the spectrum is “flat”, this suggests the planet has NO atmosphere

HOT GAS GIANT EXOPLANET WASP-39 b

ATMOSPHERE COMPOSITION

NIRSpec PRISM

*Transit
Depth*



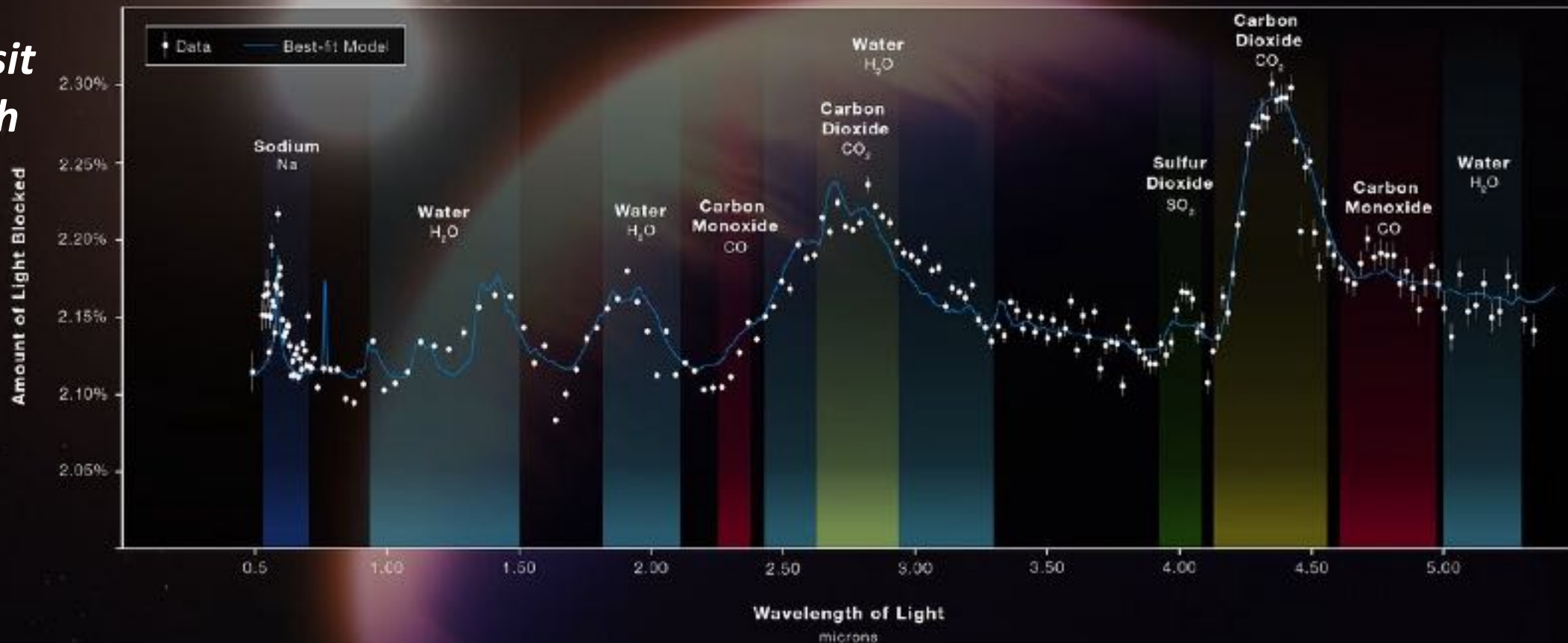
WEBB
SPACE TELESCOPE

Credit: NASA, ESA, CSA, Joseph Olmsted (STScI)

HOT GAS GIANT EXOPLANET WASP-39 b ATMOSPHERE COMPOSITION

NIRSpec PRISM

*Transit
Depth*



WEBB
SPACE TELESCOPE

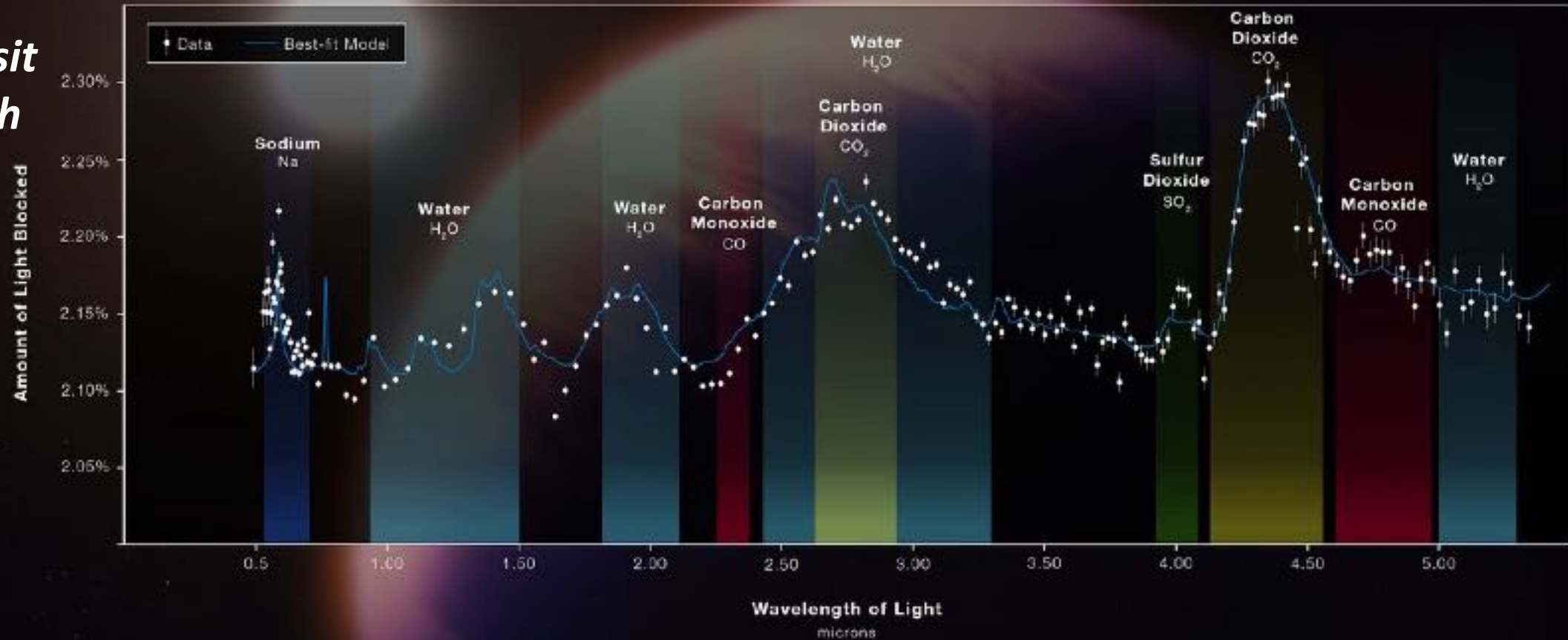
Credit: NASA, ESA, CSA, Joseph Olmsted (STScI)

HOT GAS GIANT EXOPLANET WASP-39 b ATMOSPHERE COMPOSITION

Transmission spectrum

NIRSpec PRISM

*Transit
Depth*



WEBB
SPACE TELESCOPE

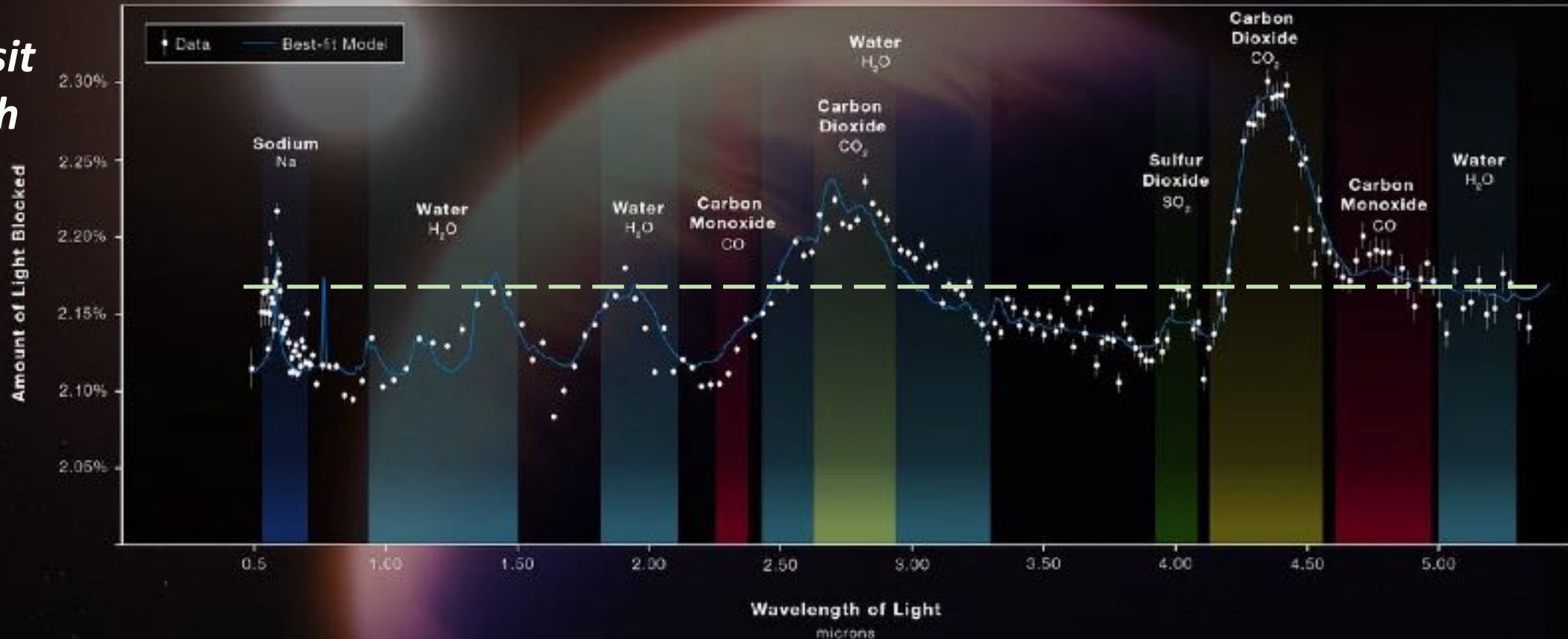
Credit: NASA, ESA, CSA, Joseph Olmsted (STScI)

HOT GAS GIANT EXOPLANET WASP-39 b
ATMOSPHERE COMPOSITION

Transmission spectrum

NIRSpec PRISM

*Transit
Depth*



FLAT LINE for rocky planets means NO
ATMOSPHERE (most likely)

WEBB
SPACE TELESCOPE

Credit: NASA, ESA, CSA, Joseph Olmsted (STScI)

ROCKY EXOPLANET LHS 475 b

TRANSMISSION SPECTRUM

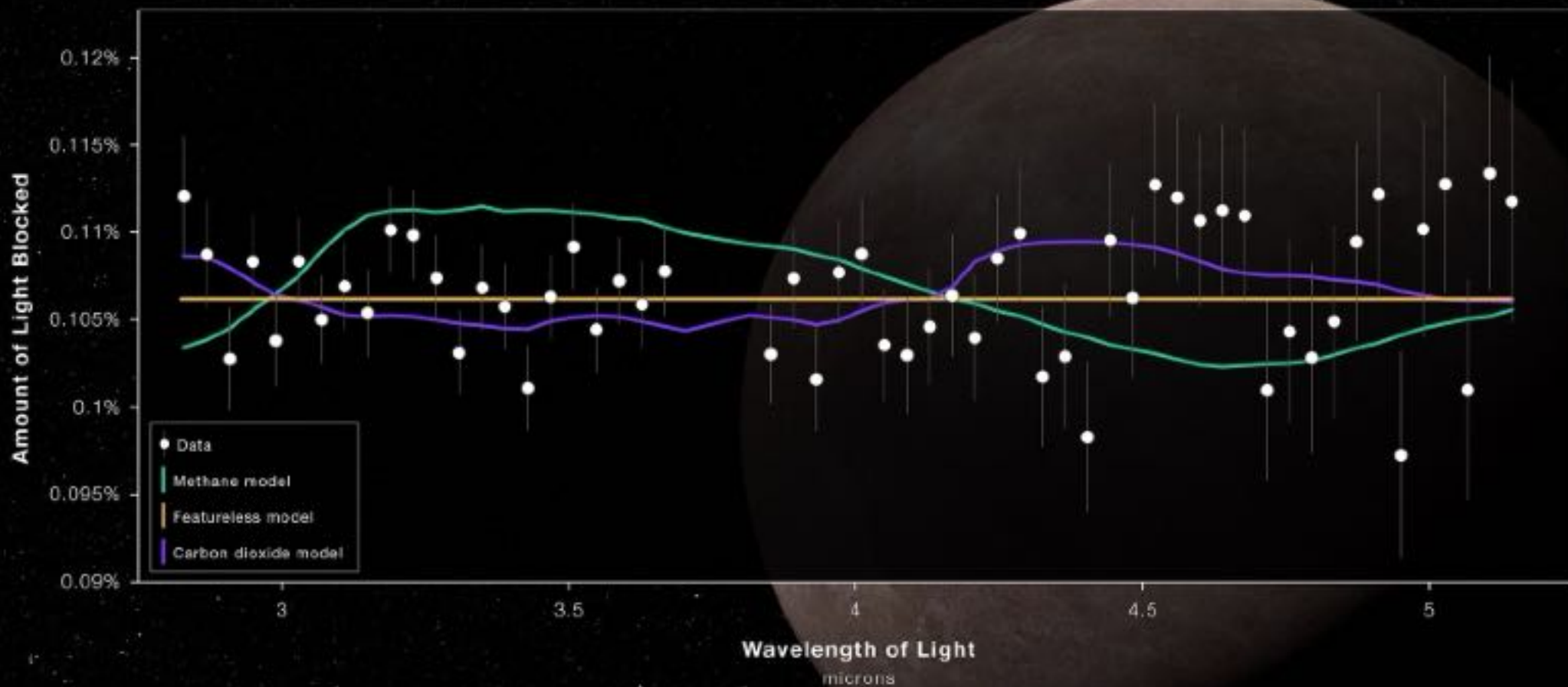
NIRSpec | Bright Object Time-Series Spectroscopy



ROCKY EXOPLANET LHS 475 b

TRANSMISSION SPECTRUM

NIRSpec | Bright Object Time-Series Spectroscopy



K2-18b

K2-18b

An Earth-like planet? Tantalising proof of life on THIS exoplanet has scientists in 'shock'

Cambridge, England • Edited By: Trisha Pathak • Updated: Apr 29, 2024, 03:08 PM IST



Forget Mars, are there aliens on... K2-18b? Discovery of planet twice as big as Earth emitting gas 'only produced by life' sparks huge excitement among astronomers

Planet K2-18b: 5 Facts About "Super Earth" Being Probed For Alien Life

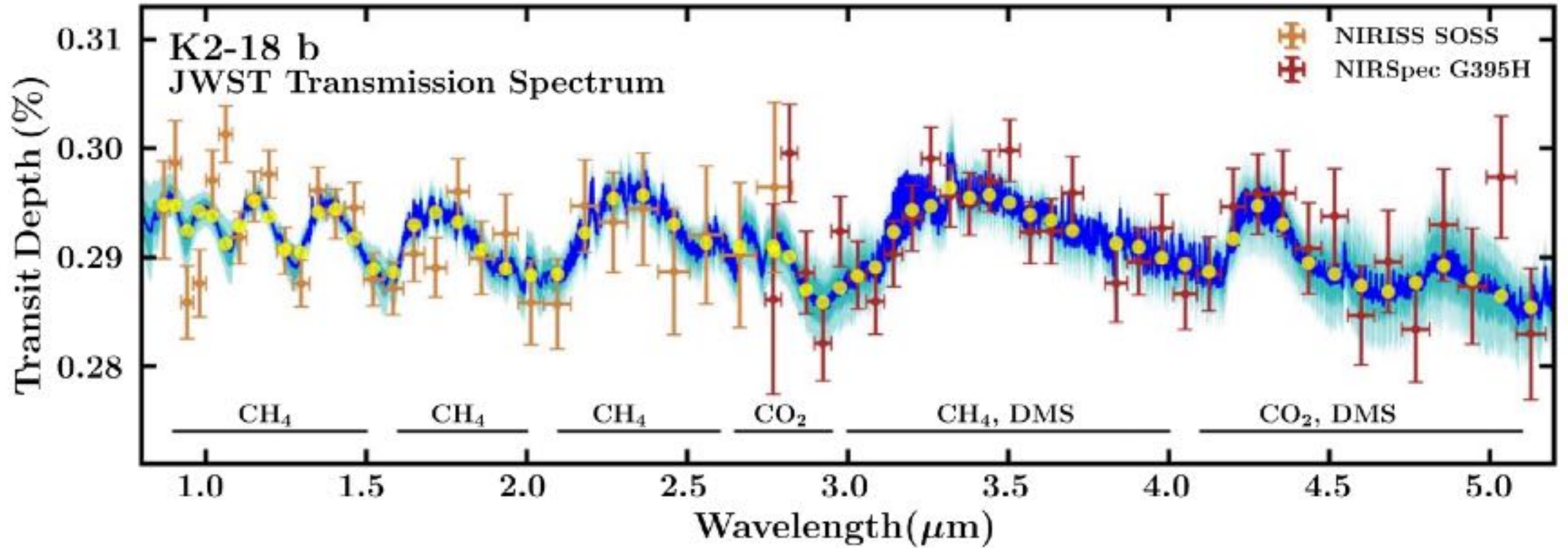
K2-18b orbits a red dwarf star, which is smaller and dimmer than our sun, and it completes one orbit in 32.9 days.

Science | Edited by NDTV News Desk | Updated: April 29, 2024 12:21 pm IST

K2-18b

**Webb Discovers Methane,
Carbon Dioxide in
Atmosphere of K2-18 b**

K2-18b





Right. Where are we going?



